Three-Dimensional Sound Field Reproduction Based on Weighted Mode-Matching Method

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Abstract-A sound field reproduction method based on the spherical wavefunction expansion of sound fields is proposed, which can be flexibly applied to various array geometries and directivities. First, we formulate sound field synthesis as a minimization problem of some norm on the difference between the desired and synthesized sound fields, and then the optimal driving signals are derived by using the spherical wavefunction expansion of the sound fields. This formulation is closely related to the mode-matching method; a major advantage of the proposed method is the optimal weight on the mode determined according to the norm to be minimized instead of the empirical truncation in the mode-matching method. We also provide some examples of norms and their corresponding weights in analytical forms. Both interior and exterior sound field reproduction are considered in the proposed method, and some applications, such as multizone reproduction and interior reproduction with exterior cancellation, are also discussed. Numerical simulation results indicated that higher reproduction accuracy can be achieved by the proposed method than by the current pressure-matching and mode-matching methods.

Index Terms—Mode-matching method, sound field reproduction, spherical wavefunction expansion.

I. INTRODUCTION

T HE aim of sound field reproduction is to physically synthesize a desired sound field over a target region using multiple loudspeakers. Owing to its potential application to high-fidelity audio systems, it has attracted much attention in the field of spatial-audio research in recent years.

One major approach is an analytic method based on the assumption of a continuous loudspeaker distribution. Furthermore, this approach is classified into two types according to its underlying theory: Kirchhoff–Helmholtz or Rayleigh integrals [1]–[5] (e.g., Wave Field Synthesis) and the spatial Fourier representation of sound fields [6]–[9] (e.g., Spectral Division Method). In these methods, the desired sound field is reproduced over the entire space enclosed by the loudspeaker array or over the half space bounded by a planar array in the case of Wave Field

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Synthesis. However, these methods basically require a simple array geometry such as a line, plane, circle, or sphere, which obviously limits their scope of practical application.

Several methods such as the pressure-matching method [10]-[12] and mode-matching method [3], [13]–[15] can be more flexibly applied to various array configurations and directivities. The pressure-matching method is formulated so that the synthesized sound pressures correspond to the desired ones at the given discrete control points, and the driving signals are usually obtained by the least-squares method with regularization. Owing to its simple formulation, this method can be applied in almost any situation; however, inevitable errors occur away from the control points, especially in cases when the control points are coarsely arranged. Although this problem can be prevented to some extent by finely arranging the control points over the target region, the computational cost also increases with the number of control points. On the other hand, the mode-matching method aims to match the modes of the synthesized and desired sound fields at a certain control point, which is often used in Higher-Order Ambisonics approach. A "mode" in the three-dimensional (3D) context usually means a spherical wavefunction, by which a sound field can be expanded, and the expansion coefficient is also called an ambisonic component or coefficient in the literature of Higher-Order Ambisonics [16]. Since the spherical wavefunctions, which are the products of spherical Bessel functions and spherical harmonics, are the solutions of the Helmholtz equation, the mode-matching method exploits its constraint. This is a major difference between the mode-matching method and the pressure-matching method. However, it is necessary in the mode-matching method to truncate the modes in an empirical manner, which strongly affects its reproduction accuracy. Note that an excessively large or small truncation order leads to performance degradation in the mode-matching method.

For the two-dimensional (2D) case, Betlehem and Abhayapala [15] proposed the weighting of expansion coefficients in the mode-matching method to minimize the spatial reproduction error in the context of reverberant sound field control. Our previous works [17]–[19] addressed the weights of expansion coefficients, and their advantage over the empirical truncation in the mode-matching method was shown for the 2D case. Since the optimal weightings on the expansion coefficients are analytically derived according to the objective function to be minimized (e.g., the spatial squared error of the sound pressure inside the given target region), these methods can be referred to as *weighted mode-matching methods*. Only specific cases in the 2D space

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were considered in these works; therefore, more general cases in the 3D space have to be investigated.

In this study, our goal is twofold: to establish the general concept of the weighted mode-matching method in the 3D space and to propose several examples of weights that are derived in analytical forms. We consider both interior and exterior sound field reproduction, where sound fields are expanded by the interior and exterior spherical wavefunctions, respectively.

The rest of this paper is organized as follows. In Section II, we introduce several notations and basic theories on the spherical wavefunction expansion of sound fields, which are used throughout this paper. In Section III, two current methods, the pressure-matching method and the mode-matching method, are briefly introduced. In Section IV, we describe the proposed method, where both the general concept of the weighted mode-matching method and several examples of the norms and corresponding weights are given. We also discuss some possible applications such as multizone reproduction [20]–[22] and interior reproduction with exterior cancellation [5], [13], [19]. In Section V, several numerical experiments performed to evaluate the proposed method are described and their results are given. Finally, we present our conclusions in Section VI.

II. NOTATIONS AND PRELIMINARIES

First, we introduce several notations and basic theories used throughout this paper.

A. Notations

Suppose that L secondary sources, i.e., loudspeakers, are located at r_1, \ldots, r_L with an arbitrary geometry in the 3D space \mathbb{R}^3 . The *l*th loudspeaker's driving signal at (angular) frequency ω is denoted by $d_l(\omega)$, and the vector of the driving signals is denoted by $d(\omega) = [d_1(\omega), \ldots, d_L(\omega)]^T$, where the superscript T denotes the transpose. Similarly, the *l*th loudspeaker's transfer function at position r and frequency ω is denoted by $g_l(r, \omega)$. Then, the synthesized sound pressure $u_{syn}(r, \omega)$ is represented as

$$u_{\rm syn}(\boldsymbol{r},\omega) = \sum_{l=1}^{L} d_l(\omega) g_l(\boldsymbol{r},\omega). \tag{1}$$

Hereafter, we omit ω for notational simplicity. In addition, we refer to the function $u_{\rm syn}$ as the synthesized sound field, and we also use the same notation for all other sound fields.

We address two sound field reproduction problems: interior sound field reproduction and exterior sound field reproduction. Let S_{int} be a spherical surface centered at r_c not including any secondary sources and S_{ext} be a spherical surface centered at r_c including all the secondary sources as shown in Fig. 1. We define Ω_{int} as a region inside S_{int} and Ω_{ext} as a region outside S_{ext} . An interior sound field is defined as a sound field such that the sound sources exist only outside S_{int} , and an exterior sound field is defined as a sound field such that the sound sources exist only inside S_{ext} . The objective of interior sound field reproduction is to reproduce a given desired interior sound field inside S_{int} , and that of exterior sound field reproduction is to reproduce a

Fig. 1. Settings of secondary sources and target region. (a) Interior sound field reproduction; (b) exterior sound field reproduction.

given desired exterior sound field outside S_{ext} . In either case, the desired sound field is denoted by u_{des} .

B. Spherical Wavefunction Expansion of Interior Sound Fields

In the case of interior sound field reproduction, g_l and u_{des} can be expanded at the expansion center r_c as [23]

$$u_{\rm des}(\boldsymbol{r}) = \sum_{\nu,\mu} \mathring{u}_{\rm des,\nu,\mu}^{\rm int}(\boldsymbol{r}_{\rm c})\varphi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\rm c}), \qquad (2)$$

$$g_l(\boldsymbol{r}) = \sum_{\nu,\mu} \mathring{g}_{l,\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}})\varphi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\text{c}}), \qquad (3)$$

where $\sum_{\nu,\mu}$ is the abbreviated form of $\sum_{\nu=0}^{\infty} \sum_{\mu=-\nu}^{\nu}$. In (2) and (3), $\mathring{u}_{\text{des},\nu,\mu}^{\text{int}}(\mathbf{r}_{\text{c}})$ and $\mathring{g}_{l,\nu,\mu}^{\text{int}}(\mathbf{r}_{\text{c}})$ are the interior expansion coefficients at the expansion center \mathbf{r}_{c} , and $\varphi_{\nu,\mu}(\cdot)$ is the interior spherical wavefunction defined as

$$\varphi_{\nu,\mu}(\boldsymbol{r}) = \sqrt{4\pi j_{\nu}(kr)}Y_{\nu,\mu}(\hat{\boldsymbol{r}}),\tag{4}$$

where $k = \omega/c$ is the wavenumber at the sound speed $c, j_{\nu}(\cdot)$ is the ν th-order spherical Bessel function of the first kind [24], $Y_{\nu,\mu}(\cdot)$ is the spherical harmonic function of order ν and degree μ [24], and r and \hat{r} are defined respectively as $r = ||\mathbf{r}||$ and $\hat{\mathbf{r}} = \mathbf{r}/r$ with the Euclidean norm $||\cdot||$.

The interior expansion coefficients of the transfer functions and the desired sound field can be obtained in several ways. If they are modeled or represented analytically (e.g., by multipoles or plane waves), their interior expansion coefficients can be calculated using several mathematical formulae [3], [24]. Another approach is to directly measure the interior expansion coefficients by using a microphone array [3], [25]–[31]. Hereafter, $\mathring{u}_{\text{des},\nu,\mu}^{\text{int}}(\mathbf{r}_{\text{c}})$ and $\mathring{g}_{l,\nu,\mu}^{\text{int}}(\mathbf{r}_{\text{c}})$ are assumed to be given.

C. Spherical Wavefunction Expansion of Exterior Sound Fields

In the case of exterior sound field reproduction, g_l and u_{des} can be expanded at the expansion center r_c as [23], [32]

$$u_{\rm des}(\boldsymbol{r}) = \sum_{\nu,\mu} \mathring{u}_{\rm des,\nu,\mu}^{\rm ext}(\boldsymbol{r}_{\rm c})\psi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\rm c}),\tag{5}$$

$$g_l(\boldsymbol{r}) = \sum_{\nu,\mu} \mathring{g}_{l,\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}})\psi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\text{c}}).$$
(6)



In (5) and (6), $\dot{u}_{\text{des},\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}})$ and $\dot{g}_{l,\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}})$ are the exterior expansion coefficients at the expansion center $\boldsymbol{r}_{\text{c}}$, and $\psi_{\nu,\mu}(\cdot)$ is the exterior spherical wavefunction defined as

$$\psi_{\nu,\mu}(\boldsymbol{r}) = \sqrt{4\pi} h_{\nu}(kr) Y_{\nu,\mu}(\hat{\boldsymbol{r}}), \tag{7}$$

where $h_{\nu}(\cdot)$ is the ν th-order spherical Hankel function of the first kind [24].

Similarly to the interior expansion coefficients, these exterior expansion coefficients can be derived analytically using mathematical formulae [24] or measured by using a microphone array [29], [30]. Hereafter, $\mathring{u}_{\mathrm{des},\nu,\mu}^{\mathrm{ext}}(\boldsymbol{r}_{\mathrm{c}})$ and $\mathring{g}_{l,\nu,\mu}^{\mathrm{ext}}(\boldsymbol{r}_{\mathrm{c}})$ are assumed to be given.

D. Translation of Expansion Coefficients

When the exterior sound field u is expanded at the expansion center $r_{\rm c}$ as

$$u(\boldsymbol{r}) = \sum_{\nu,\mu} \mathring{u}_{\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}})\psi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\text{c}}), \qquad (8)$$

the expansion coefficients at a different expansion center r'_{c} are given by the following equations [24]:

$$\mathring{u}_{\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\rm c}') = \sum_{\nu',\mu'} T_{\nu,\mu'}^{\nu',\mu'}(\boldsymbol{r}_{\rm c}'-\boldsymbol{r}_{\rm c})\mathring{u}_{\nu',\mu'}^{\text{ext}}(\boldsymbol{r}_{\rm c}),\tag{9}$$

$$\mathring{u}_{\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\rm c}') = \sum_{\nu',\mu'} S_{\nu,\mu}^{\nu',\mu'}(\boldsymbol{r}_{\rm c}'-\boldsymbol{r}_{\rm c})\mathring{u}_{\nu',\mu'}^{\text{ext}}(\boldsymbol{r}_{\rm c}), \qquad (10)$$

where $T^{\nu',\mu'}_{\nu,\mu}(\cdot)$ and $S^{\nu',\mu'}_{\nu,\mu}(\cdot)$ are the translation operators defined as

$$T_{\nu,\mu}^{\nu',\mu'}(\mathbf{r}) = 4\pi (-1)^{\mu'} j^{\nu-\nu'} \cdot \sum_{s=0}^{\nu+\nu'} j^s j_s(kr) Y_{s,\mu-\mu'}(\hat{\mathbf{r}})^* \mathcal{G}(\nu',\mu';\nu,-\mu;s),$$
(11)

$$S_{\nu,\mu}^{\nu',\mu'}(\boldsymbol{r}) = 4\pi (-1)^{\mu'} j^{\nu-\nu'} \cdot \sum_{s=0}^{\nu+\nu'} j^s h_s(kr) Y_{s,\mu-\mu'}(\hat{\boldsymbol{r}})^* \mathcal{G}(\nu',\mu';\nu,-\mu;s).$$
(12)

Here, j is the imaginary unit, the superscript * denotes the complex conjugate, and $\mathcal{G}(\cdot)$ is the Gaunt coefficient [24].

III. CURRENT METHODS

We briefly introduce two current methods that are closely related to the proposed method.

A. Pressure-Matching Method

The pressure-matching method aims to control the sound pressures at given discrete control points so that they correspond to the desired ones. Regardless of the reproduction of interior or exterior sound fields, a typical formulation is as follows [11]. Let $r_{cp}^{(1)}, \ldots, r_{cp}^{(M)}$ be the positions of the control points. Then, the driving signals are derived by solving the following

minimization problem:

$$\underset{\boldsymbol{d}\in\mathbb{C}^{L}}{\text{minimize}} \quad \sum_{m=1}^{M} \left| u_{\text{syn}}(\boldsymbol{r}_{\text{cp}}^{(m)}) - u_{\text{des}}(\boldsymbol{r}_{\text{cp}}^{(m)}) \right|^{2} + \lambda \boldsymbol{d}^{\mathsf{H}}\boldsymbol{d}, \quad (13)$$

where λ is the regularization parameter and the superscript H denotes the Hermitian transpose. As derived in [11], the solution of (13) is given by

$$\hat{\boldsymbol{d}}_{\mathrm{pm}} = \left(\mathbf{A}_{\mathrm{pm}} + \lambda \boldsymbol{I}_L\right)^{-1} \mathbf{b}_{\mathrm{pm}},$$
 (14)

where I_L is the $L \times L$ identity matrix, and $\mathbf{A}_{pm} \in \mathbb{C}^{L \times L}$ and $\mathbf{b}_{pm} \in \mathbb{C}^L$ are defined as

$$(\mathbf{A}_{\rm pm})_{l_1,l_2} = \sum_{m=1}^{M} g_{l_1} (\boldsymbol{r}_{\rm cp}^{(m)})^* g_{l_2} (\boldsymbol{r}_{\rm cp}^{(m)}), \qquad (15)$$

$$(\mathbf{b}_{pm})_l = \sum_{m=1}^M g_l(\boldsymbol{r}_{cp}^{(m)})^* u_{des}(\boldsymbol{r}_{cp}^{(m)}).$$
 (16)

Here, $(\cdot)_{l_1,l_2}$ denotes the (l_1, l_2) th element of the matrix and $(\cdot)_l$ denotes the *l*th element of the vector.

In many practical systems, fixed secondary sources are used to reproduce various desired sound fields. Therefore, the computational cost after setting the desired sound field is important in these systems, whereas other values such as $A_{\rm pm}$ can be calculated beforehand. From this viewpoint, (14) can be rewritten as

$$\hat{\boldsymbol{d}}_{\mathrm{pm}} = \mathbf{C}_{\mathrm{pm}} \boldsymbol{u}_{\mathrm{des}}.$$
 (17)

where $u_{\text{des}} \in \mathbb{C}^M$ is the vector of the sound pressures at the control points and $\mathbf{C}_{\text{pm}} \in \mathbb{C}^{L \times M}$ is a matrix independent of the desired sound field. Since \mathbf{C}_{pm} can be calculated beforehand for a fixed array, the computational complexity after setting the desired sound field is $\mathcal{O}(LM)$.

B. Mode-Matching Method

The mode-matching method aims to control the expansion coefficients of the synthesized sound field so that they correspond to the desired ones. In the reproduction of interior sound fields, the driving signals are obtained from the following mode-matching equation:

$$\dot{u}_{\mathrm{syn},\nu,\mu}^{\mathrm{int}}(\boldsymbol{r}_{\mathrm{c}}) = \dot{u}_{\mathrm{des},\nu,\mu}^{\mathrm{int}}(\boldsymbol{r}_{\mathrm{c}}), \quad \forall \nu \in \{0,\dots,N_{\mathrm{int}}\}, \quad (18)$$

where N_{int} is the truncation order and $\mathring{u}_{\text{syn},\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}})$ is the interior expansion coefficient of u_{syn} given as $\mathring{u}_{\text{syn},\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}}) = \sum_{l=1}^{L} d_{l} \mathring{g}_{l,\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}})$. As shown in [3], [14], the regularized solution of (18) can be obtained by solving

$$\underset{\boldsymbol{d}\in\mathbb{C}^{L}}{\text{minimize}} \quad \sum_{\nu,\mu}^{N_{\text{int}}} \left| \mathring{u}_{\text{syn},\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}}) - \mathring{u}_{\text{des},\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}}) \right|^{2} + \lambda \boldsymbol{d}^{\mathsf{H}}\boldsymbol{d} \quad (19)$$

as

$$\hat{\boldsymbol{d}}_{\rm mm} = \left(\mathbf{A}_{\rm mm} + \lambda \boldsymbol{I}_L\right)^{-1} \mathbf{b}_{\rm mm}.$$
 (20)

Here, $\sum_{\nu,\mu}^{N_{\text{int}}}$ is the abbreviated form of $\sum_{\nu=0}^{N_{\text{int}}} \sum_{\mu=-\nu}^{\nu}$, and $\mathbf{A}_{\text{mm}} \in \mathbb{C}^{L \times L}$ and $\mathbf{b}_{\text{mm}} \in \mathbb{C}^{L}$ are given by

$$(\mathbf{A}_{\rm mm})_{l_1, l_2} = \sum_{\nu, \mu}^{N_{\rm int}} \mathring{g}_{l_1, \nu, \mu}^{\rm int}(\boldsymbol{r}_{\rm c})^* \mathring{g}_{l_2, \nu, \mu}^{\rm int}(\boldsymbol{r}_{\rm c}), \qquad (21)$$

$$(\mathbf{b}_{\mathrm{mm}})_l = \sum_{\nu,\mu}^{N_{\mathrm{int}}} \mathring{g}_{l,\nu,\mu}^{\mathrm{int}}(\boldsymbol{r}_{\mathrm{c}})^* \mathring{u}_{\mathrm{des},\nu,\mu}^{\mathrm{int}}(\boldsymbol{r}_{\mathrm{c}}).$$
(22)

Similarly, in the reproduction of exterior sound fields, the mode-matching equation is formulated as

$$\mathring{u}_{\text{syn},\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}}) = \mathring{u}_{\text{des},\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}}), \quad \forall \nu \in \{0,\dots,N_{\text{ext}}\}, \quad (23)$$

where N_{ext} is the truncation order and $\mathring{u}_{\text{syn},\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}}) = \sum_{l=1}^{L} d_{l} \mathring{g}_{l,\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}})$ is the exterior expansion coefficient of u_{syn} . Therefore, the driving signals are given by

$$\hat{\boldsymbol{d}}_{mm} = \left(\mathbf{A}_{mm} + \lambda \boldsymbol{I}_L \right)^{-1} \mathbf{b}_{mm},$$
 (24)

where $\mathbf{A}_{\mathrm{mm}} \in \mathbb{C}^{L \times L}$ and $\mathbf{b}_{\mathrm{mm}} \in \mathbb{C}^{L}$ are defined as

$$(\mathbf{A}_{\rm mm})_{l_1, l_2} = \sum_{\nu, \mu}^{N_{\rm ext}} \mathring{g}_{l_1, \nu, \mu}^{\rm ext}(\boldsymbol{r}_{\rm c})^* \mathring{g}_{l_2, \nu, \mu}^{\rm ext}(\boldsymbol{r}_{\rm c}), \qquad (25)$$

$$(\mathbf{b}_{\rm mm})_l = \sum_{\nu,\mu}^{N_{\rm ext}} \mathring{g}_{l,\nu,\mu}^{\rm ext}(\boldsymbol{r}_{\rm c})^* \mathring{u}_{{\rm des},\nu,\mu}^{\rm ext}(\boldsymbol{r}_{\rm c}).$$
(26)

The truncation orders N_{int} and N_{ext} are usually determined empirically according to the size of the array of the secondary sources or the reproduction region. However, it is difficult to find the optimal truncation orders, especially in interior sound field reproduction; an excessively large N_{int} degrades the reproduction accuracy, which can also be seen in the experiments in Section V.

Either (20) or (24) can be rewritten as

$$\hat{\boldsymbol{d}}_{\mathrm{mm}} = \mathbf{C}_{\mathrm{mm}} \hat{\boldsymbol{u}}_{\mathrm{des}},$$
 (27)

where $\hat{\boldsymbol{u}}_{\mathrm{des}} \in \mathbb{C}^{\check{N}}$ is the vector of the interior or exterior expansion coefficients of the desired sound field and $\mathbf{C}_{\mathrm{mm}} \in \mathbb{C}^{L \times \check{N}}$ is a matrix independent of the desired sound field. Here, \check{N} is defined as $\check{N} = (N_{\mathrm{int}} + 1)^2$ in the case of interior sound field reproduction and as $\check{N} = (N_{\mathrm{ext}} + 1)^2$ in the case of exterior sound field reproduction. Therefore, the computational complexity after setting the desired sound field is $\mathcal{O}(L\check{N})$.

IV. PROPOSED METHOD

In Sections IV-A and Section IV-B, the general concept of the proposed weighted mode-matching method is described. Also, we provide several examples of weights on expansion coefficients in Sections IV-C and Section IV-D. Finally, we discuss the application of the proposed method to multizone reproduction and interior sound field reproduction with exterior cancellation in Section IV-E.

A. Formulation

The sound field reproduction problem can be formulated as minimization of the difference between the desired and synthesized sound fields. Therefore, we consider the following general formulation:

$$\underset{\boldsymbol{d}\in\mathbb{C}^{L}}{\operatorname{minimize}} \quad J(\boldsymbol{d}) = \|\boldsymbol{u}_{\operatorname{syn}} - \boldsymbol{u}_{\operatorname{des}}\|_{(*)}^{2} + \lambda \boldsymbol{d}^{\mathsf{H}}\boldsymbol{d}, \qquad (28)$$

where $\|\cdot\|_{(*)}$ is some norm on interior or exterior sound fields induced by some inner product $\langle\cdot,\cdot\rangle_{(*)}$ as

$$||u||_{(*)}^2 = \langle u, u \rangle_{(*)} .$$
⁽²⁹⁾

A typical example of this type of norm is the L_2 norm with any positive spatial weight. Therefore, one can set various formulations according to the detailed objective. Several examples of norms are described in Section IV-C for interior sound fields and in Section IV-D for exterior sound fields. From (1) and (29), The objective function of (28) can be rewritten as

$$J(\boldsymbol{d}) = \boldsymbol{d}^{\mathsf{H}}(\mathbf{A}_{\mathrm{wmm}} + \lambda \boldsymbol{I}_{L})\boldsymbol{d} - \boldsymbol{d}^{\mathsf{H}}\mathbf{b}_{\mathrm{wmm}} - \mathbf{b}_{\mathrm{wmm}}^{\mathsf{H}}\boldsymbol{d} + \mathbf{c}_{\mathrm{wmm}},$$
(30)

where $\mathbf{A}_{\text{wmm}} \in \mathbb{C}^{L \times L}$, $\mathbf{b}_{\text{wmm}} \in \mathbb{C}^{L}$, and $\mathbf{c}_{\text{wmm}} \in [0, \infty)$ are given by

$$(\mathbf{A}_{\text{wmm}})_{l_1, l_2} = \langle g_{l_1}, g_{l_2} \rangle_{(*)},$$
 (31)

$$(\mathbf{b}_{\mathrm{wmm}})_l = \langle g_l, u_{\mathrm{des}} \rangle_{(*)}, \tag{32}$$

$$c_{\rm wmm} = \langle u_{\rm des}, u_{\rm des} \rangle_{(*)}. \tag{33}$$

Since $\mathbf{A}_{\text{wmm}} + \lambda \mathbf{I}_L$ is a positive definite Hermitian matrix, i.e., $\mathbf{x}^{\mathsf{H}}(\mathbf{A}_{\text{wmm}} + \lambda \mathbf{I}_L)\mathbf{x} > 0$ for any $\mathbf{x} \in \mathbb{C}^L \setminus \{\mathbf{0}\}$, it has an inverse (and also positive definite Hermitian) matrix $(\mathbf{A}_{\text{wmm}} + \lambda \mathbf{I}_L)^{-1}$. Therefore, J can be represented as

$$J(d) = \left(d - (\mathbf{A}_{wmm} + \lambda \mathbf{I}_L)^{-1} \mathbf{b}_{wmm} \right)^{\mathsf{H}} \\ \cdot \left(\mathbf{A}_{wmm} + \lambda \mathbf{I}_L \right) \left(d - (\mathbf{A}_{wmm} + \lambda \mathbf{I}_L)^{-1} \mathbf{b}_{wmm} \right) \\ + c_{wmm} - \mathbf{b}_{wmm}^{\mathsf{H}} (\mathbf{A}_{wmm} + \lambda \mathbf{I}_L)^{-1} \mathbf{b}_{wmm}, \quad (34)$$

and the optimal driving signals $\hat{d}_{\text{wmm}} \in \mathbb{C}^{L}$, which minimize J, are obtained as

$$\hat{\boldsymbol{d}}_{\text{wmm}} = \left(\mathbf{A}_{\text{wmm}} + \lambda \boldsymbol{I}_L \right)^{-1} \mathbf{b}_{\text{wmm}}.$$
 (35)

This can also be derived on the basis of the Wirtinger derivative, i.e., by solving $\frac{\partial}{\partial d^{\text{H}}} J(d) = 0$. Therefore, the remaining problem is the calculation of \mathbf{A}_{wmm} and \mathbf{b}_{wmm} .

B. Weighted Mode-Matching

Consider an inner product of two interior sound fields u_1 and u_2 , i.e., $\langle u_1, u_2 \rangle_{(*)}$. Let $\mathring{u}_{1,\nu,\mu}(\mathbf{r}_c)$ and $\mathring{u}_{2,\nu,\mu}(\mathbf{r}_c)$ denote respectively the interior expansion coefficients of u_1 and u_2 . Then, $\langle u_1, u_2 \rangle_{(*)}$ can be rewritten by assuming the interchangeability

of infinite summation and inner product¹ as

$$\langle u_1, u_2 \rangle_{(*)} = \left\langle \sum_{\nu, \mu} \mathring{u}_{1,\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}}) \varphi_{\nu,\mu}(\boldsymbol{r} - \boldsymbol{r}_{c}), \right.$$

$$\left. \sum_{\nu, \mu} \mathring{u}_{2,\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}}) \varphi_{\nu,\mu}(\boldsymbol{r} - \boldsymbol{r}_{c}) \right\rangle_{(*)}$$

$$= \left. \sum_{\nu_{1}, \mu_{1}} \sum_{\nu_{2}, \mu_{2}} \mathring{u}_{1,\nu_{1},\mu_{1}}^{\text{int}}(\boldsymbol{r}_{\text{c}})^{*} \mathring{u}_{2,\nu_{2},\mu_{2}}^{\text{int}}(\boldsymbol{r}_{\text{c}}) \right.$$

$$\left. \cdot \left\langle \varphi_{\nu_{1},\mu_{1}}, \varphi_{\nu_{1},\mu_{1}} \right\rangle_{(*)}.$$

$$(36)$$

Therefore, in interior sound field reproduction, ${\bf A}_{\rm wmm}$ and ${\bf b}_{\rm wmm}$ can be rewritten as

$$(\mathbf{A}_{\text{wmm}})_{l_1, l_2} = \sum_{\nu_1, \mu_1} \sum_{\nu_2, \mu_2} w_{\nu_1, \mu_1}^{\nu_2, \mu_2} \cdot \mathring{g}_{l_1, \nu_1, \mu_1}^{\text{int}}(\boldsymbol{r}_{\text{c}})^* \mathring{g}_{l_2, \nu_2, \mu_2}^{\text{int}}(\boldsymbol{r}_{\text{c}}),$$
(37)

$$(\mathbf{b}_{\text{wmm}})_{l} = \sum_{\nu_{1},\mu_{1}} \sum_{\nu_{2},\mu_{2}} w_{\nu_{1},\mu_{1}}^{\nu_{2},\mu_{2}} \cdot \mathring{g}_{l,\nu_{1},\mu_{1}}^{\text{int}}(\boldsymbol{r}_{\text{c}})^{*} \mathring{u}_{\text{des},\nu_{2},\mu_{2}}^{\text{int}}(\boldsymbol{r}_{\text{c}}),$$
(38)

where $w_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ are the weights on the interior expansion coefficients (referred to as the interior weights, hereafter) given as

$$w_{\nu_{1},\mu_{1}}^{\nu_{2},\mu_{2}} = \langle \varphi_{\nu_{1},\mu_{1}}(\boldsymbol{r}-\boldsymbol{r}_{c}), \varphi_{\nu_{2},\mu_{2}}(\boldsymbol{r}-\boldsymbol{r}_{c}) \rangle_{(*)}.$$
(39)

In exterior sound field reproduction, on the other hand, $A_{\rm wmm}$ and $b_{\rm wmm}$ can be rewritten as

$$(\mathbf{A}_{\text{wmm}})_{l_1, l_2} = \sum_{\nu_1, \mu_1} \sum_{\nu_2, \mu_2} v_{\nu_1, \mu_1}^{\nu_2, \mu_2} \cdot \mathring{g}_{l_1, \nu_1, \mu_1}^{\text{ext}}(\boldsymbol{r}_{\text{c}})^* \mathring{g}_{l_2, \nu_2, \mu_2}^{\text{ext}}(\boldsymbol{r}_{\text{c}}),$$
(40)

$$(\mathbf{b}_{\text{wmm}})_{l} = \sum_{\nu_{1},\mu_{1}} \sum_{\nu_{2},\mu_{2}} v_{\nu_{1},\mu_{1}}^{\nu_{2},\mu_{2}} \cdot \mathring{g}_{l,\nu_{1},\mu_{1}}^{\text{ext}}(\boldsymbol{r}_{\text{c}})^{*} \mathring{u}_{\text{des},\nu_{2},\mu_{2}}^{\text{ext}}(\boldsymbol{r}_{\text{c}}),$$
(41)

where $v_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ are the weights on the exterior expansion coefficients (referred to as the exterior weights, hereafter) given as

$$v_{\nu_1,\mu_1}^{\nu_2,\mu_2} = \langle \psi_{\nu_1,\mu_1}(\boldsymbol{r} - \boldsymbol{r}_{\rm c}), \psi_{\nu_2,\mu_2}(\boldsymbol{r} - \boldsymbol{r}_{\rm c}) \rangle_{(*)}.$$
(42)

Therefore, we only have to calculate the inner product on the spherical wavefunctions, i.e., $w_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ and $v_{\nu_1,\mu_1}^{\nu_2,\mu_2}$, and \mathbf{A}_{wmm} and \mathbf{b}_{wmm} can be calculated for arbitrary g_l and u_{des} using (37) and (38) or (40) and (41). Note that $w_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ and $v_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ are independent of g_l and u_{des} ; therefore, we can calculate them beforehand analytically or numerically. Since these infinite summations are difficult to solve in many cases, the summations are truncated at a sufficiently large order N_{int} or N_{ext} . Comparing (37), (38), (40), and (41) with (21), (22), (25), and (26), respectively, one can see that the difference between the proposed method and the

current mode-matching method appears in the weights on the expansion coefficients, i.e., $w_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ and $v_{\nu_1,\mu_1}^{\nu_2,\mu_2}$. In the proposed method, the optimal weights on the expansion coefficients are derived according to the norm to be minimized.

We also evaluate the computational cost after the desired sound field is given. Equation (35) can be rewritten in a similar form to (27) as

$$d_{\rm wmm} = \mathbf{C}_{\rm wmm} \dot{\boldsymbol{u}}_{\rm des},$$
 (43)

where $\mathbf{C}_{\text{wmm}} \in \mathbb{C}^{L \times \check{N}}$ is independent of the desired sound field. Therefore, the computational complexity after setting the desired sound field is $\mathcal{O}(L\check{N})$, which is the same as that of the mode-matching method.

C. Examples of Norms on Interior Sound Fields and Their Interior Weights

We here give two examples of norms on interior sound fields and their corresponding interior weights, which are also given in Table I.

1) Interior Uniformly Weighted L_2 Norm: The first example is the interior uniformly weighted L_2 norm $\|\cdot\|_{int,uni}$, which is defined as

$$\|u\|_{\text{int,uni}}^2 = \int_{\Omega_{\text{int}}} |u(\boldsymbol{r})|^2 \,\mathrm{d}\boldsymbol{r}.$$
 (44)

By using this norm, one can reproduce the sound field within the spherical region Ω_{int} . In this case, the interior weights are obtained analytically as

$$w_{\nu_1,\mu_1}^{\nu_2,\mu_2} = \delta_{\nu_1,\nu_2} \delta_{\mu_1,\mu_2} w_{\mathrm{uni},\nu_1}, \tag{45}$$

where $\delta_{n,m}$ denotes the Kronecker's delta and

$$w_{\mathrm{uni},\nu} = 2\pi R_{\mathrm{int}}^3 \left(j_{\nu} (kR_{\mathrm{int}})^2 - j_{\nu-1} (kR_{\mathrm{int}}) j_{\nu+1} (kR_{\mathrm{int}}) \right), \quad (46)$$

where R_{int} is the radius of the sphere S_{int} . The proof is given

in Appendix B.

2) Interior Gaussian-Weighted L_2 Norm: The second example is the interior Gaussian-weighted L_2 norm $\|\cdot\|_{int,Gs}$, which is defined as

$$\|u\|_{\text{int,Gs}}^2 = \int_{\Omega_{\text{int}}} \exp\left(-\frac{\|\boldsymbol{r} - \boldsymbol{r}_c\|^2}{2\sigma^2}\right) |u(\boldsymbol{r})|^2 \,\mathrm{d}\boldsymbol{r}, \quad (47)$$

where σ is the scale parameter. Note that the Gaussian w(r) is defined only on Ω_{int} and truncated outside it. By using this norm, one can prioritize the reproduction around the central region while preventing significant deterioration in the peripheral region. In this case, the interior weights are obtained as

$$w_{\nu_1,\mu_1}^{\nu_2,\mu_2} = \delta_{\nu_1,\nu_2} \delta_{\mu_1,\mu_2} w_{\mathrm{Gs},\nu_1}, \tag{48}$$

where

$$w_{\text{Gs},\nu} = 4\pi \int_{0}^{R_{\text{int}}} \exp\left(-\frac{r^{2}}{2\sigma^{2}}\right) j_{\nu}(kr)^{2} \,\mathrm{d}r,$$

$$= \sum_{n=0}^{\infty} (-1)^{n} 2\pi R_{\text{int}}^{3} \exp\left(-\frac{R_{\text{int}}^{2}}{4\sigma^{2}}\right) i_{n}\left(\frac{R_{\text{int}}^{2}}{4\sigma^{2}}\right)$$

$$\cdot (j_{\nu-n}(kR_{\text{int}})j_{\nu+n}(kR_{\text{int}}))$$

$$- j_{\nu-n-1}(kR_{\text{int}})j_{\nu+n+1}(kR_{\text{int}})), \qquad (49)$$

¹The interchangeability of infinite summation and inner product holds if the spherical wavefunction expansion of sound fields converges with respect to the norm $\|\cdot\|_{(*)}$ (see Appendix A for the proof). However, this condition is satisfied in almost any practical case (at least when the norm $\|u\|_{(*)}$ consists of the integral of u and its partial derivatives on a compact set) including all examples shown in Sections IV-C and Section IV-D. This is because the spherical wavefunction expansion of the interior sound field and its partial derivatives converge uniformly in Ω_{int} and the spherical wavefunction expansion of the exterior sound field and its partial derivatives converge uniformly in any compact subset of Ω_{ext} [23].

 TABLE I

 Examples of Norms on Interior Sound Fields and Their Interior Weights

Norm on interior sound fields	Interior weights
Interior uniformly weighted L_2 norm:	
$\ u\ _{ ext{int,uni}}^2 = \int_{\Omega_{ ext{int}}} u(oldsymbol{r}) ^2 \mathrm{d}oldsymbol{r}$	$w_{\text{uni},\nu} = 2\pi R_{\text{int}}^3 \left(j_{\nu} (kR_{\text{int}})^2 - j_{\nu-1} (kR_{\text{int}}) j_{\nu+1} (kR_{\text{int}}) \right)$
Interior Gaussian-weighted L_2 norm: $\ u\ _{\text{int,Gs}}^2 = \int_{\Omega_{\text{int}}} \exp\left(-\frac{\ \boldsymbol{r}-\boldsymbol{r}_{\text{c}}\ ^2}{2\sigma^2}\right) u(\boldsymbol{r}) ^2 \mathrm{d}\boldsymbol{r}$	$ w_{\text{Gs},\nu} = \sum_{n=0}^{\infty} (-1)^n 2\pi R_{\text{int}}^3 \exp\left(-\frac{R_{\text{int}}^2}{4\sigma^2}\right) i_n\left(\frac{R_{\text{int}}^2}{4\sigma^2}\right) \cdot (j_{\nu-n}(kR_{\text{int}})j_{\nu+n}(kR_{\text{int}}) - j_{\nu-n-1}(kR_{\text{int}})j_{\nu+n+1}(kR_{\text{int}})) $

 TABLE II

 Examples of Norms on Exterior Sound Fields and Their Exterior Weights

Norm on exterior sound fields	Exterior weights
Exterior uniformly weighted L_2 norm: $\ u\ _{\text{ext,uni}}^2 = \int_{\Omega'_{\text{ext}}} u(\mathbf{r}) ^2 \mathrm{d}\mathbf{r}$	$v_{\mathrm{uni},\nu} = 2\pi R_{\mathrm{ext}}^{\prime 3} \left(h_{\nu}(kR_{\mathrm{ext}}) ^2 - h_{\nu-1}(kR_{\mathrm{ext}})^* h_{\nu+1}(kR_{\mathrm{ext}}) \right) - 2\pi R_{\mathrm{ext}}^3 \left(h_{\nu}(kR_{\mathrm{ext}}) ^2 - h_{\nu-1}(kR_{\mathrm{ext}})^* h_{\nu+1}(kR_{\mathrm{ext}}) \right)$
Exterior radiation-power norm: $\ u\ _{\text{ext,rad}}^2 = \int_{S_{\text{ext}}} \frac{1}{2} \operatorname{Re} \left[u(\boldsymbol{r}) \frac{j}{\rho c k} \frac{\partial}{\partial \hat{\boldsymbol{n}}(\boldsymbol{r})} u(\boldsymbol{r})^* \right] d\boldsymbol{r}$	$v_{\rm rad,\nu} = \frac{2\pi}{\rho c k^2}$



Fig. 2. Example of interior weights for interior uniformly weighted L_2 norm (Uni) and interior Gaussian-weighted L_2 norm (Gs).

where $i_n(\cdot)$ is the *n*th-order modified spherical Bessel function of the first kind [24]. The derivation is given in Appendix C. Although there is no useful formula to evaluate the integral in the first line of (49), it can be calculated efficiently by truncating the second line at sufficiently large *n*.

In both examples, the orthogonality of the interior spherical wavefunctions holds as in (45) and (48). Therefore, (37) and (38) can be rewritten as

$$(\mathbf{A}_{\text{wmm}})_{l_1, l_2} = \sum_{\nu, \mu} w_{(*), \nu} \mathring{g}_{l_1, \nu, \mu}^{\text{int}}(\mathbf{r}_{\text{c}})^* \mathring{g}_{l_2, \nu, \mu}^{\text{int}}(\mathbf{r}_{\text{c}}), \qquad (50)$$

$$(\mathbf{b}_{\text{wmm}})_l = \sum_{\nu,\mu} w_{(*),\nu} \mathring{g}_{l,\nu,\mu}^{\text{int}}(\mathbf{r}_{\text{c}})^* \mathring{u}_{\text{des},\nu,\mu}^{\text{int}}(\mathbf{r}_{\text{c}}),$$
 (51)

where $w_{(*),\nu}$ represents either $w_{\mathrm{uni},\nu}$ or $w_{\mathrm{Gs},\nu}$. As an example, the values of $w_{\mathrm{uni},\nu}$ or $w_{\mathrm{Gs},\nu}$ are plotted in Fig. 2 in the case of k = 10 rad/m, $R_{\mathrm{int}} = 1.2 \text{ m}$, and $\sigma = 0.6 \text{ m}$. Here, the weights are normalized so that their zeroth-order values correspond to one.

D. Examples of Norms on Exterior Sound Fields and Their Exterior Weights

We also give two examples of norms on exterior sound fields and their exterior weights, which are also given in Table II.

1) Exterior Uniformly Weighted L_2 Norm: The first example is the exterior uniformly weighted L_2 norm $\|\cdot\|_{ext,uni}$, which is defined as

$$\|u\|_{\text{ext,uni}}^2 = \int_{\Omega_{\text{ext}}} |u(\boldsymbol{r})|^2 \,\mathrm{d}\boldsymbol{r}.$$
 (52)

Here, $\Omega'_{\text{ext}} \subset \Omega_{\text{ext}}$ is defined as the region between the inner sphere S_{ext} and the outer sphere S'_{ext} , which is also centered at r_{c} . By using this norm, one can reproduce the sound field within the region Ω'_{ext} . In this case, the exterior weights are obtained analytically as

$$\nu_{\nu_1,\mu_1}^{\nu_2,\mu_2} = \delta_{\nu_1,\nu_2} \delta_{\mu_1,\mu_2} v_{\mathrm{uni},\nu_1}, \tag{53}$$

where

$$v_{\mathrm{uni},\nu} = 2\pi R_{\mathrm{ext}}^{\prime 3} \left(|h_{\nu} (kR_{\mathrm{ext}})|^2 - h_{\nu-1} (kR_{\mathrm{ext}})^* h_{\nu+1} (kR_{\mathrm{ext}}) \right) - 2\pi R_{\mathrm{ext}}^3 \left(|h_{\nu} (kR_{\mathrm{ext}})|^2 - h_{\nu-1} (kR_{\mathrm{ext}})^* h_{\nu+1} (kR_{\mathrm{ext}}) \right).$$
(54)

Here, R_{ext} and R'_{ext} are the radii of S_{ext} and S'_{ext} , respectively. The derivation is given in Appendix D.

2) Exterior Radiation-Power Norm: The second example is the exterior radiation-power norm $\|\cdot\|_{ext,rad}$, which is defined as

$$\|u\|_{\text{ext,rad}}^{2} = \int_{S_{\text{ext}}} \frac{1}{2} \operatorname{Re} \left[u(\boldsymbol{r}) \frac{\mathrm{j}}{\rho c k} \frac{\partial}{\partial \hat{\boldsymbol{n}}(\boldsymbol{r})} u(\boldsymbol{r})^{*} \right] \mathrm{d}\boldsymbol{r}, \quad (55)$$

where Re[·] denotes the real part of a complex value, ρ is the density of air, and $\partial/\partial \hat{n}(r)$ denotes the outward normal derivative. Since the integrand of (55) represents the outward acoustic intensity [33], this norm corresponds to the total power



Fig. 3. Example of exterior weights for exterior uniformly weighted L_2 norm (Uni) and exterior radiation-power norm (Rad).

radiated outside S_{ext} . In this case, the exterior weights are obtained analytically as

$$v_{\nu_1,\mu_1}^{\nu_2,\mu_2} = \delta_{\nu_1,\nu_2} \delta_{\mu_1,\mu_2} v_{\mathrm{rad},\nu_1}, \tag{56}$$

where

$$v_{\mathrm{rad},\nu} = \frac{2\pi}{\rho c k^2}.$$
(57)

The derivation is given in Appendix E.

In both examples, the orthogonality of the exterior spherical wavefunctions holds as in (53) and (56). Therefore, (40) and (41) can be rewritten as

$$(\mathbf{A}_{\text{wmm}})_{l_1, l_2} = \sum_{\nu, \mu} v_{(*), \nu} \mathring{g}_{l_1, \nu, \mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}})^* \mathring{g}_{l_2, \nu, \mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}}), \qquad (58)$$

$$(\mathbf{b}_{\mathrm{wmm}})_{l} = \sum_{\nu,\mu} v_{(*),\nu} \mathring{g}_{l,\nu,\mu}^{\mathrm{ext}}(\boldsymbol{r}_{\mathrm{c}})^{*} \mathring{u}_{\mathrm{des},\nu,\mu}^{\mathrm{ext}}(\boldsymbol{r}_{\mathrm{c}}), \qquad (59)$$

where $v_{(*),\nu}$ represents either $v_{\mathrm{uni},\nu}$ or $v_{\mathrm{rad},\nu}$. As an example, the values of $v_{\mathrm{uni},\nu}$ or $v_{\mathrm{rad},\nu}$ are plotted in Fig. 3 in the case of k = 10 rad/m, $R_{\mathrm{ext}} = 2.0$ m, and $R'_{\mathrm{ext}} = 2.5$ m. Here, the weights are normalized so that their zeroth-order values correspond to one.

Furthermore, in the case of the exterior radiation-power norm, (58) and (59) can be calculated analytically without infinite summations as follows. Suppose each g_l and u_{des} are modeled by the finite-order exterior spherical wavefunctions around their source positions r_l and r_{des} (e.g., monopole, dipole, and cardioid sources) as

$$g_l(\boldsymbol{r}) = \sum_{\nu,\mu}^{N_l} \beta_{l,\nu,\mu} \psi_{\nu,\mu}(\boldsymbol{r} - \boldsymbol{r}_l), \qquad (60)$$

$$u_{\rm des}(\boldsymbol{r}) = \sum_{\nu,\mu}^{N_{\rm des}} \beta_{{\rm des},\nu,\mu} \psi_{\nu,\mu}(\boldsymbol{r} - \boldsymbol{r}_{\rm des}). \tag{61}$$

In this case, $\mathbf{A}_{\mathrm{wmm}}$ and $\mathbf{b}_{\mathrm{wmm}}$ are derived analytically as

$$(\mathbf{A}_{\text{wmm}})_{l_1, l_2} = \frac{2\pi}{\rho c k^2} \sum_{\nu_1, \mu_1}^{N_{l_1}} \sum_{\nu_2, \mu_2}^{N_{l_2}} T^{\nu_2, \mu_2}_{\nu_1, \mu_1} (\boldsymbol{r}_{l_1} - \boldsymbol{r}_{l_2})$$

$$\beta^*_{l_1, \nu_1, \mu_1} \beta_{l_2, \nu_2, \mu_2}, \qquad (62)$$

$$(\mathbf{b}_{\text{wmm}})_{l} = \frac{2\pi}{\rho c k^{2}} \sum_{\nu_{1},\mu_{1}}^{N_{l}} \sum_{\nu_{2},\mu_{2}}^{N_{\text{des}}} T_{\nu_{1},\mu_{1}}^{\nu_{2},\mu_{2}} (\boldsymbol{r}_{l} - \boldsymbol{r}_{\text{des}}) \cdot \beta_{l,\nu_{1},\mu_{1}}^{*} \beta_{\text{des},\nu_{2},\mu_{2}}.$$
(63)

The derivation is given in Appendix E. Note that \mathbf{A}_{wmm} and \mathbf{b}_{wmm} do not depend on \mathbf{r}_{c} or R_{ext} in this case; they are determined only by the relative positions and directivities of the secondary and desired sources.

E. Several Applications

In the above sections, only a single norm is used to formulate the minimization problem; however, we can formulate various minimization problems by combining multiple norms. Although there are various possible applications, we here discuss two examples: multizone reproduction and interior reproduction with exterior cancellation.

1) Multizone Reproduction: Multizone reproduction aims to reproduce different sound fields in different listening areas. Let $q = 1, \ldots, Q$ be the index of the listening area, $\Omega_{int}^{(q)}$ be the *q*th listening area, which is a spherical region centered at $r_c^{(q)}$, and $u_{des}^{(q)}$ denote the desired interior sound field in the *q*th listening area. Then, the multizone reproduction is formulated as

$$\underset{\boldsymbol{d}\in\mathbb{C}^{L}}{\operatorname{minimize}} \quad J(\boldsymbol{d}) = \sum_{q=1}^{Q} \gamma_{q} \|\boldsymbol{u}_{\operatorname{syn}} - \boldsymbol{u}_{\operatorname{des}}^{(q)}\|_{(q)}^{2} + \lambda \boldsymbol{d}^{\mathsf{H}}\boldsymbol{d}, \quad (64)$$

where $\|\cdot\|_{(q)}$ is the norm on the interior sound fields defined on $\Omega_{\text{int}}^{(q)}$ and γ_q is a constant parameter used to balance the Qnorms. In this case, the optimal driving signals are obtained as

$$\hat{\boldsymbol{d}} = \left(\sum_{q=1}^{Q} \gamma_q \mathbf{A}^{(q)} + \lambda \boldsymbol{I}_L\right)^{-1} \sum_{q=1}^{Q} \gamma_q \mathbf{b}^{(q)}, \qquad (65)$$

where $\mathbf{A}^{(q)}$ and $\mathbf{b}^{(q)}$ are derived for each q with a different expansion center $\mathbf{r}_{c}^{(q)}$.

2) Interior Reproduction With Exterior Cancellation: In the reproduction of interior sound fields in reverberant environments, it is useful to suppress the exterior radiation outside the secondary sources to prevent the performance degradation caused by reverberation. This can be formulated as follows:

$$\underset{\boldsymbol{d} \in \mathbb{C}^{L}}{\operatorname{minimize}} \quad J(\boldsymbol{d}) = \|\boldsymbol{u}_{\operatorname{syn}} - \boldsymbol{u}_{\operatorname{des}}\|_{\operatorname{int},(*)}^{2}$$
$$+ \eta \|\boldsymbol{u}_{\operatorname{syn}}\|_{\operatorname{ext},(*)}^{2} + \lambda \boldsymbol{d}^{\mathsf{H}} \boldsymbol{d}, \qquad (66)$$

where $\|\cdot\|_{int,(*)}$ is some norm on interior sound fields, $\|\cdot\|_{ext,(*)}$ is some norm on exterior sound fields, and η is a constant parameter used to balance the two norms. In this case, the optimal driving signals are obtained as

$$\hat{\boldsymbol{d}} = (\mathbf{A}_{\text{int}} + \eta \mathbf{A}_{\text{ext}} + \lambda \boldsymbol{I}_L)^{-1} \mathbf{b}_{\text{int}},$$
 (67)

where \mathbf{A}_{int} and \mathbf{b}_{int} are respectively given by (37) and (38) on the basis of the norm $\|\cdot\|_{int,(*)}$ and \mathbf{A}_{ext} is given by (40) on the basis of the norm $\|\cdot\|_{ext,(*)}$.

V. NUMERICAL EXPERIMENTS

Numerical simulations were conducted to evaluate the proposed method in comparison with the pressure-matching method [11] and the mode-matching method [13]. Hereafter, the proposed method, the pressure-matching method, and the mode-matching method are denoted by **Proposed**, **PM**, and **MM**, respectively.

A. Interior Sound Field Reproduction Using Cardioid Loudspeaker Array

In Cartesian coordinates r = (x, y, z), 144 secondary sources were located on a sphere with a radius of 1.5 m centered at (0, 0, 0) m.

The positions of the secondary sources were determined on the basis of the spherical *t*-design [34]. Each secondary source was modeled as a fixed-directivity first-order source, whose transfer function is given by [14]

$$g_{l}(\boldsymbol{r}) = \frac{\exp(jk\|\boldsymbol{r} - \boldsymbol{r}_{l}\|)}{4\pi\|\boldsymbol{r} - \boldsymbol{r}_{l}\|} \cdot \left[\alpha - (1-\alpha)\left(1 + \frac{j}{k\|\boldsymbol{r} - \boldsymbol{r}_{l}\|}\right)\cos\gamma_{l}\right], \quad (68)$$

where α is a directivity parameter and γ_l represents the angle between $\mathbf{r} - \mathbf{r}_l$ and the highest-gain direction of the secondary source, denoted by $\hat{\mathbf{p}}_l$. The directivity parameters were set as $\alpha = 0.5$ (i.e., cardioid source) and the highest-gain directions are oriented to the center of the sphere, i.e., $\hat{\mathbf{p}}_l = -\hat{\mathbf{r}}_l$, where $\hat{\mathbf{r}}_l = \mathbf{r}_l / ||\mathbf{r}_l||$. Then, g_l can be represented as [14]

$$g_l(\boldsymbol{r}) = \sum_{\nu,\mu} \beta_{l,\nu,\mu} \psi_{\nu,\mu}(\boldsymbol{r} - \boldsymbol{r}_l)$$
(69)

with

$$\beta_{l,\nu,\mu} = \begin{cases} \alpha \frac{jk}{\sqrt{4\pi}}, & \nu = 0\\ (1-\alpha)kY_{1,\mu}(-\hat{\boldsymbol{p}}_l)^*, & \nu = 1.\\ 0, & \nu \ge 2 \end{cases}$$
(70)

The desired sound field was a plane wave propagating in the positive x-direction, i.e., $u_{\text{des}}(r) = a \exp(jkx)$, where the amplitude was set as a = 1. The sound speed was set as 340.29 m/s.

In **Proposed** and **MM**, the center of the spherical wavefunction expansion was set at $r_c = (0, 0, 0)$ m. The interior expansion coefficients of the secondary sources were calculated using (70) and (12), and the interior expansion coefficients of the desired sound field were calculated as [3]

$$\mathring{u}_{\mathrm{des},\nu,\mu}^{\mathrm{int}}(\boldsymbol{r}_{\mathrm{c}}) = a\sqrt{4\pi} \mathsf{j}^{\nu} Y_{\nu,\mu}(\hat{\boldsymbol{r}}_{x})^{*},\tag{71}$$

where \hat{r}_x is the unit direction vector along the positive *x*direction. Several truncation orders $N_{\rm int}$ were used in **Proposed** and **MM**. In **Proposed**, the two norms described in Section IV-C, i.e., $\|\cdot\|_{\rm int,uni}$ and $\|\cdot\|_{\rm int,Gs}$, were evaluated, which are hereafter denoted by **Proposed (Uni)** and **Proposed** (**Gs**), respectively. The radius of $S_{\rm int}$ was set as $R_{\rm int} = 1.2$ m, and the scale parameter σ in **Proposed (Gs)** was set as 0.3 m. In **PM**, the control points were aligned in a grid at regular



Fig. 4. Normalized reproduction error (NRE) plotted against number of basis functions at 550 Hz when plane wave propagating in positive *x*-direction is reproduced by using spherical array of 144 cardioid sources oriented inward.

intervals in all three directions within $\Omega_{\rm int}$, where various intervals were evaluated. The regularization parameter λ was set as $\lambda = 10^{-3} \times ||\mathbf{A}||_2$ in all the methods, where $||\mathbf{A}||_2$ denotes the maximum singular value of \mathbf{A} and \mathbf{A} represents $\mathbf{A}_{\rm wmm}$, $\mathbf{A}_{\rm mm}$, and $\mathbf{A}_{\rm pm}$ for **Proposed**, **MM**, and **PM**, respectively.

To evaluate the reproduction accuracy, we define the normalized reproduction error (NRE) as

$$\operatorname{NRE}(\omega) = 10 \log_{10} \frac{\int_{\Omega_{\operatorname{int}}} |u_{\operatorname{syn}}(\boldsymbol{r},\omega) - u_{\operatorname{des}}(\boldsymbol{r},\omega)|^2 \,\mathrm{d}\boldsymbol{r}}{\int_{\Omega_{\operatorname{int}}} |u_{\operatorname{des}}(\boldsymbol{r},\omega)|^2 \,\mathrm{d}\boldsymbol{r}} \tag{dB},$$
(72)

where the spatial integrals were discretized with an interval of 0.05 m. Note that small NRE indicates high reproduction performance.

First, we plotted the relationship of the reproduction accuracy to the number of basis functions, i.e., \tilde{N} for **Proposed** and MM and M for PM, at 550 Hz in Fig. 4. Proposed and MM achieved almost the same NREs for a small number of basis functions. However, the NRE for MM rapidly increased when the number of basis functions exceeded 169, i.e., $N_{\rm int} > 12$. This is considered to be because the same weightings are imposed on all the modes. From the properties of the spherical Bessel functions of the first kind [3], spherical wavefunctions of high orders tend to have small values within Ω_{int} . When large N_{int} is used, a relatively large effort is devoted to matching these insignificant modes, and the significant low-order modes are undervalued in the MM, which causes the increase of the NREs. Note that the optimal $N_{int} = 12$ for **MM** does not correspond to either $\lceil kR_{\text{int}} \rceil = 10$ or $\lceil (e/2)kR_{\text{int}} \rceil = 14$, which are two commonly used rules of thumb on the truncation [3], [21]. Therefore, it can be seen that one has to evaluate many different $N_{\rm int}$ in **MM** to achieve the best performance. On the other hand, **Proposed** is not affected by N_{int} as long as it is sufficiently large. This is an important point because it implies that the tuning of Nrequired in MM is unnecessary in Proposed, which can result in less total complexity of computation. Although the NREs for Proposed (Uni) and PM were close for a large number of basis functions, that for **Proposed(Uni)** reached its peak at a relatively smaller number of basis functions than that for PM. This means



Fig. 5. Reproduced sound pressure distributions at 550 Hz when plane wave propagating in positive x-direction is reproduced using spherical array of 144 cardioid sources oriented inward. (a) **Proposed (Uni)**; (b) **Proposed (Gs)**; (c) **MM**; (d) **PM**.



Fig. 6. Normalized error distributions at 550 Hz when plane wave propagating in positive x-direction is reproduced using spherical array of 144 cardioid sources oriented inward. (a) **Proposed (Uni)**; (b) **Proposed (Gs)**; (c) **MM**; (d) **PM**.

that a lower computational cost can be achieved for **Proposed** than that for **PM**.

Figs. 5 and 6 respectively show the reproduced sound pressure distributions and the normalized error distributions at 550 Hz. The solid lines and dashed lines represent the sphere where secondary sources exist and the reproduction region Ω_{int} , respectively. In addition, we plotted NREs against distance from the center of the array defined as

NRE
$$(r,\omega)$$

= $10 \log_{10} \frac{\int_{\mathbb{S}^2} |u_{\text{des}}(r\hat{\boldsymbol{r}},\omega) - u_{\text{syn}}(r\hat{\boldsymbol{r}},\omega)|^2 \,\mathrm{d}\hat{\boldsymbol{r}}}{\int_{\mathbb{S}^2} |u_{\text{des}}(r\hat{\boldsymbol{r}},\omega)|^2 \,\mathrm{d}\hat{\boldsymbol{r}}} (\text{dB})$ (73)

in Fig. 7. Here, \mathbb{S}^2 denotes the unit sphere and the integral was calculated by Monte Carlo integration using 500 points sampled from the spherical uniform distribution. On the basis of the result in Fig. 4, truncation orders were determined here as $N_{\text{int}} = 12$ in **Proposed** and **MM**. In **PM**, the interval of the control points was set as 0.35 m so that the number of control points slightly exceeded the number of basis functions of the other methods, i.e., M = 175 and $\tilde{N} = 169$. The NREs of **Proposed (Uni)**, **Proposed (Gs)**, **MM**, and **PM** were -13.16, -12.08, -11.56, and -0.26 dB, respectively. **Proposed (Uni)** achieved the lowest NRE, and its reproduction accuracy was relatively uniformly high within Ω_{int} . On the other hand, **Proposed(Gs)** achieved much higher reproduction accuracy around the central region of Ω_{int} than the other methods. This is due to the interior weights based on the norm $\|\cdot\|_{Gs}$.



Fig. 7. Normalized reproduction error (NRE) plotted against distance from center at 550 Hz when plane wave propagating in positive x-direction is reproduced by using spherical array of 144 cardioid sources oriented inward.

The NRE is also plotted against frequency from 50 Hz to 800 Hz at intervals of 50 Hz in Fig. 8. Since it will take a high computational cost to seek the optimal N_{int} at each frequency, especially in **MM**, two rules of thumb, $N_{\text{int}} = \lceil kR_{\text{int}} \rceil$ and $N_{\text{int}} = \lceil (e/2)kR_{\text{int}} \rceil$, were investigated in **MM**. In **Proposed**, N_{int} was set as $\lceil (e/2)kR_{\text{int}} \rceil$, and the number of control points in **PM** was determined so that it was equal to or slightly larger than \breve{N} for **Proposed**. **Proposed** (**Gs**) was omitted here since it does not aim the uniformly high reproduction accuracy within Ω_{int} . Again, it can be seen the performance of **MM** depends heavily on the truncation order. Note that superiority and inferiority between two rules differ by frequencies, which implies



Fig. 8. Normalized reproduction error (NRE) plotted against frequency when plane wave propagating in positive *x*-direction is reproduced by using spherical array of 144 cardioid sources oriented inward.

the complexity of tuning of N_{int} in **MM**. Without such tuning of N_{int} , **Proposed (Uni)** achieved the lowest NRE at all the frequencies. This is due to the optimal weighting on the interior expansion coefficients determined analytically on the basis of the norm on the sound fields.

B. Exterior Sound Field Reproduction Using Cardioid Loudspeaker Array

We also conducted a simulation of exterior sound field reproduction. In this experiment, 144 cardioid sources were located at the same position as in Section V-A, and each cardioid source was oriented outward from the center of the array, i.e, $\hat{p}_l = \hat{r}_l$. The desired sound field was a spherical wave from a point source at $r_{\text{des}} = (1.0, 0, 0)$ m, i.e., $u_{\text{des}}(r) = a \exp(jk||r - r_{\text{des}}||)/4\pi ||r - r_{\text{des}}||$, where the amplitude was set as a = 10.

In **Proposed** and **MM**, the center of the spherical wavefunction expansion was set at $r_c = (0, 0, 0)$ m. The exterior expansion coefficients of the secondary sources were calculated using (70) and (11), and the exterior expansion coefficients of the desired sound field were calculated as [3]

$$\mathring{u}_{\mathrm{des},\nu,\mu}^{\mathrm{ext}}(\boldsymbol{r}_{\mathrm{c}}) = a \frac{\mathrm{j}k}{\sqrt{4\pi}} j_{\nu}(kr_{\mathrm{des}})Y_{\nu,\mu}(\hat{\boldsymbol{r}}_{\mathrm{des}})^{*},$$
 (74)

where $r_{des} = ||\mathbf{r}_{des}||$ and $\hat{\mathbf{r}}_{des} = \mathbf{r}_{des}/r_{des}$. Several truncation orders N_{ext} were used in **Proposed** and **MM**. In **Proposed**, the two norms described in Section IV-D, i.e., $|| \cdot ||_{ext,uni}$ and $|| \cdot ||_{ext,rad}$, were evaluated, which are hereafter denoted by **Proposed (Uni)** and **Proposed (Rad)**, respectively. In **Proposed** (**Uni**), the radii of S_{ext} and S'_{ext} were respectively set as $R_{ext} =$ 2.0 m and $R_{ext} = 2.5$ m. In **Proposed (Rad)**, \mathbf{A}_{wmm} and \mathbf{b}_{wmm} were calculated analytically using (62) and (63). In **PM**, the control points were aligned in a grid at regular intervals in all three directions within Ω'_{ext} , where various intervals were evaluated. The regularization parameter λ was set as $\lambda = 10^{-3} \times ||\mathbf{A}||_2$ in all the methods. To evaluate the reproduction accuracy, we



Fig. 9. Nomalized reproduction error (NRE) plotted against number of basis functions at 400 Hz when spherical wave from point source at (1.0, 0, 0) m is reproduced by using spherical array of 144 cardioid sources oriented outward.

define the NRE as

$$\operatorname{NRE}(\omega) = 10 \log_{10} \frac{\int_{\Omega_{\text{ext}}'} |u_{\text{syn}}(\boldsymbol{r},\omega) - u_{\text{des}}(\boldsymbol{r},\omega)|^2 \,\mathrm{d}\boldsymbol{r}}{\int_{\Omega_{\text{ext}}'} |u_{\text{des}}(\boldsymbol{r},\omega)|^2 \,\mathrm{d}\boldsymbol{r}} \tag{dB},$$
(75)

where the spatial integrals were discretized with an interval of 0.05 m.

First, we plotted the relationship of the reproduction accuracy to the number of basis functions at 400 Hz in Fig. 9. **Proposed** (**Rad**) was omitted because no truncation was used in it. In this case, the NREs for **Proposed** (**Uni**) and **MM** were almost the same, and the NRE for **MM** did not increase as N_{ext} increased. Note that **MM** corresponds to **Proposed** (**Rad**) as $N_{\text{ext}} \rightarrow \infty$. As in the interior sound field reproduction, **Proposed** and **MM** achieved their maximum performance at a smaller number of basis functions than **PM**.

Figs. 10 and 11 respectively show the reproduced sound pressure distributions and the normalized error distributions at 400 Hz. The dashed lines represent the reproduction region Ω'_{ext} . On the basis of the result in Fig. 9, the truncation order was determined as $N_{ext} = 13$ in **Proposed (Uni)** and **MM**. In **PM**, the interval of control points was set as 0.55 m so that the number of control points slightly exceeded the number of basis functions of the other methods, i.e., M = 204 and $\breve{N} = 196$. The NREs for **Proposed (Uni)**, **Proposed (Rad)**, **MM**, and **PM** were -17.43, -17.45, -17.40, and -15.12 dB. The NREs for all the methods were close; however, **Proposed (Rad)** achieved the lowest value.

The NRE is also plotted against frequency from 50 Hz to 800 Hz at intervals of 50 Hz in Fig. 12. Here, the truncation order was set as $N_{\text{ext}} = \lceil (e/2)kR_{\text{ext}} \rceil$ in **Proposed (Uni)** and **MM**. The number of control points in **PM** was determined so that it was equal to or slightly larger than \breve{N} for the other methods. It can be seen that **Proposed (Uni)** achieved higher performance than the other methods, especially at low frequencies.

C. Multizone Reproduction With Exterior Cancellation Using Omnidirectional Loudspeaker Array

Finally, we evaluate **Proposed** for multizone reproduction with exterior cancellation using a monopole array. The



Fig. 10. Reproduced sound pressure distributions at 400 Hz when spherical wave from point source at (1.0, 0, 0) m is reproduced by using spherical array of 144 cardioid sources oriented outward. (a) **Proposed (Uni)**; (b) **Proposed (Rad)**; (c) **MM**; (d) **PM**.



Fig. 11. Normalized error distributions at 400 Hz when spherical wave from point source at (1.0, 0, 0) m is reproduced by using spherical array of 144 cardioid sources oriented outward. (a) **Proposed (Uni)**; (b) **Proposed (Rad)**; (c) **MM**; (d) **PM**.



Fig. 12. Normalized reproduction error (NRE) plotted against frequency when spherical wave from point source at (1.0, 0, 0) m is reproduced by using spherical array of 144 cardioid sources oriented outward.

monopole array was configured as follows. First, in the xyplane, 64 monopole sources were aligned equidistantly on double squares centered at (0,0,0) m with an inner side length of 3.0 m and an outer side length of 3.4 m. Then, five sets of these 64 monopole sources were aligned equidistantly along the z-direction at intervals of 0.5 m. Therefore, the total number of secondary sources was L = 320. The number of multizones was set as Q = 2, and the center and radius of $\Omega_{\rm int}^{(1)}$ and $\Omega_{\rm int}^{(2)}$ were $r_{\rm c}^{(1)} = (0, 0.8, 0)$ m, $R_{\rm int}^{(1)} = 0.4$ m and $r_{\rm c}^{(2)} = (0, -0.8, 0)$ m, $R_{\rm int}^{(2)} = 0.4$ m, respectively. The desired multizone sound fields were defined as $u_{\rm des}^{(1)}(r) = a \exp(jkx)$ on $\Omega_{\rm int}^{(1)}$ with a = 1 and $u_{\rm des}^{(2)}(\mathbf{r}) = 0$ on $\Omega_{\rm int}^{(2)}$. In addition, exterior cancellation was attempted; therefore, the driving signals were obtained as

$$\hat{\boldsymbol{d}} = \left(\mathbf{A}_{\text{int}}^{(1)} + \mathbf{A}_{\text{int}}^{(2)} + \eta \mathbf{A}_{\text{ext}} + \lambda \boldsymbol{I}_L\right)^{-1} \mathbf{b}_{\text{int}}^{(1)}, \quad (76)$$

where $\mathbf{A}_{\mathrm{int}}^{(1)}$, $\mathbf{A}_{\mathrm{int}}^{(2)}$, and $\mathbf{b}_{\mathrm{int}}^{(1)}$ were calculated on the basis of the interior uniformly weighted L_2 norm and $\mathbf{A}_{\mathrm{ext}}$ was calculated on the basis of the exterior radiation-power norm. The truncation orders were set as $N_{\mathrm{int}}^{(1)} = \lceil (e/2)kR_{\mathrm{int}}^{(1)} \rceil$ for $\mathbf{A}_{\mathrm{int}}^{(1)}$ and $\mathbf{b}_{\mathrm{int}}^{(1)}$ and $N_{\mathrm{int}}^{(2)} = \lceil (e/2)kR_{\mathrm{int}}^{(2)} \rceil$ for $\mathbf{A}_{\mathrm{int}}^{(2)}$. The parameters η and λ were determined experimentally as $\eta = 10^{-2} \times \rho ck^2/2\pi$ and $\lambda = 10^{-3} \times \|\mathbf{A}_{\mathrm{int}}^{(1)} + \mathbf{A}_{\mathrm{int}}^{(2)} + \eta \mathbf{A}_{\mathrm{ext}}\|_2$. Fig. 13 shows the reproduction results at 400 Hz.

Fig. 13 shows the reproduction results at 400 Hz. The black dots represent the positions of the monopole sources. Here, the normalized error distribution is calculated as $10 \log_{10} |u_{\rm syn}(\mathbf{r}) - u_{\rm des}^{(1)}(\mathbf{r})|^2 / |u_{\rm des}^{(1)}(\mathbf{r})|^2$ (dB) and the normalized power distribution is calculated as $10 \log_{10} |u_{\rm syn}(\mathbf{r})|^2 / |u_{\rm des}^{(1)}(\mathbf{r})|^2$ (dB). It can be seen that the normalized error is below -30 dB at almost all the points within $\Omega_{\rm int}^{(1)}$ and that the normalized power is below -30 dB at almost all the secondary sources.

In addition, to compare the performance of **Proposed** with **MM**, we also evaluate **MM** by using (21), (22), and (25) for $\mathbf{A}_{\text{int}}^{(1)}$, $\mathbf{A}_{\text{int}}^{(2)}$, $\mathbf{b}_{\text{int}}^{(1)}$, and \mathbf{A}_{ext} . In **MM**, the truncation order was set as $N_{\text{int}} = [\xi k R_{\text{int}}]$ and $N_{\text{ext}} = [\xi k R_{\text{ext}}]$, where $R_{\text{ext}} =$



Fig. 13. Results of multizone reproduction with exterior cancellation at 400 Hz for **Proposed**. (a) Reproduced sound pressure distribution; (b) normalized error distribution; (c) normalized power distribution.



Fig. 14. Relationship between normalized reproduction error (NRE) and normalized radiation power (NRP) in multizone reproduction with exterior cancellation at 400 Hz.

 $\max \{ \|\boldsymbol{r}_l\| \}_{l=1}^{L} = 2.53 \text{ m, and we investigated } \xi = 1 \text{ and } \xi = e/2 \text{ on the basis of the two rules of thumb. The regularization parameter in$ **MM**is determined in the same way as in**Proposed** $, i.e., <math>\lambda = 10^{-3} \times \|\mathbf{A}_{\text{int}}^{(1)} + \mathbf{A}_{\text{int}}^{(2)} + \eta \mathbf{A}_{\text{ext}} \|_2.$

The NRE and normalized radiation power (NRP) were used as evaluation criteria:

NRE = 10 log₁₀
$$\frac{\sum_{q=1}^{2} \int_{\Omega_{\text{int}}^{(q)}} |u_{\text{syn}}(\boldsymbol{r}) - u_{\text{des}}^{(q)}(\boldsymbol{r})|^2 \,\mathrm{d}\boldsymbol{r}}{\sum_{q=1}^{2} \int_{\Omega_{\text{int}}^{(q)}} |u_{\text{des}}^{(q)}(\boldsymbol{r})|^2 \,\mathrm{d}\boldsymbol{r}} \tag{dB},$$

$$NRP = 10 \log_{10} \frac{\int_{\Omega'_{ext}} |u_{syn}(\boldsymbol{r})|^2 \,\mathrm{d}\boldsymbol{r}}{\int_{\Omega'_{ext}} |a|^2 \,\mathrm{d}\boldsymbol{r}} \ (dB),$$
(78)

where Ω'_{ext} was set as the region between the inner sphere with a radius of 3.0 m and the outer sphere with a radius of 3.5 m that are both centered at (0, 0, 0) m, and the spatial integrals were discretized with an interval of 0.05 m. Here, NRP is normalized by a instead of u_{des} because $u_{\text{des}}(\mathbf{r}) = 0$ in Ω_{ext} .

Since NRE and NRP have a trade-off relationship balanced by η , we investigated many different η and plotted the relationship of NRE and NRP at 400 Hz in Fig. 14. Lower-left points indicate high reproduction accuracy with little radiation outside the secondary sources. It can be seen **Proposed** achieved lower NRE and NRP than **MM**.

VI. CONCLUSION

We proposed a sound field reproduction method based on the spherical wavefunction expansion of sound fields. Instead of the empirical truncation in the current mode-matching method, the optimal weights on the expansion coefficients are derived analytically on the basis of the norm to be minimized in the proposed method. We also showed that the proposed method is easily applicable to multizone sound field reproduction and interior sound field reproduction with exterior cancellation. Numerical experiments confirmed that higher performance with lower computational cost can be achieved by the proposed method than by the mode-matching and pressure-matching methods.

Appendix A

SUFFICIENT CONDITION OF INTERCHANGEABILITY OF INFINITE SUMMATION AND INNER PRODUCT

In (36), the interchangeability of infinite summation and an inner product holds if the spherical wavefunction expansion of sound fields converges with respect to the norm $\|\cdot\|_{(*)}$. More particularly, the norm $\|\cdot\|_{(*)}$ has to satisfy the following condition:

$$\lim_{N \to \infty} \|\rho_N\|_{(*)} = 0, \tag{79}$$

where ρ_N is the function defined as

$$\rho_N(\boldsymbol{r}) = u(\boldsymbol{r}) - \sum_{\nu,\mu}^N \mathring{u}_{\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}})\varphi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\text{c}}) \qquad (80)$$

for any interior sound field u in interior sound field reproduction and

$$\rho_N(\boldsymbol{r}) = u(\boldsymbol{r}) - \sum_{\nu,\mu}^N \mathring{u}_{\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{\text{c}})\psi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\text{c}})$$
(81)

for any exterior sound field u in exterior sound field reproduction. Here, we provide the proof. First, let $\tilde{u}_{1,N}$ and $\tilde{u}_{2,N}$ be defined as

$$\tilde{u}_{1,N}(\boldsymbol{r}) = \sum_{\nu,\mu}^{N} \mathring{u}_{1,\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}})\varphi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\text{c}}), \qquad (82)$$

$$\tilde{u}_{2,N}(\boldsymbol{r}) = \sum_{\nu,\mu}^{N} \mathring{u}_{2,\nu,\mu}^{\text{int}}(\boldsymbol{r}_{\text{c}})\varphi_{\nu,\mu}(\boldsymbol{r}-\boldsymbol{r}_{\text{c}}).$$
(83)

We also define $\rho_{1,N}$ and $\rho_{1,2}$ as $\rho_{1,N} = u_1 - \tilde{u}_{1,N}$ and $\rho_{2,N} = u_2 - \tilde{u}_{2,N}$, respectively. Then, since a finite summation and an inner product can be interchanged, $\langle u_1, u_2 \rangle_{(*)}$ can be represented as the following form for any natural numbers N_1 and N_2 :

$$\langle u_{1}, u_{2} \rangle_{(*)} = \langle \tilde{u}_{1,N_{1}} + \rho_{1,N_{1}}, \tilde{u}_{2,N_{2}} + \rho_{1,N_{2}} \rangle_{(*)}$$

$$= \langle \tilde{u}_{1,N_{1}}, \tilde{u}_{2,N_{1}} \rangle_{(*)}$$

$$+ \langle \tilde{u}_{1,N_{1}}, \rho_{2,N_{2}} \rangle_{(*)} + \langle \rho_{1,N_{2}}, \tilde{u}_{2,N_{2}} \rangle_{(*)}$$

$$= \sum_{\nu_{1},\mu_{1}}^{N_{1}} \sum_{\nu_{2},\mu_{2}}^{N_{2}} w_{\nu_{1},\mu_{1}}^{\nu_{2},\mu_{2}} \cdot \mathring{u}_{1,\nu_{1},\mu_{1}}^{\text{int}}(\boldsymbol{r}_{c})^{*} \mathring{u}_{2,\nu_{2},\mu_{2}}^{\text{int}}(\boldsymbol{r}_{c})$$

$$+ \langle u_{1}, \rho_{2,N_{2}} \rangle_{(*)} + \langle \rho_{1,N_{1}}, u_{2} \rangle_{(*)}$$

$$- \langle \rho_{1,N_{1}}, \rho_{2,N_{2}} \rangle_{(*)}.$$

$$(84)$$

Here, the following inequality holds from the Cauchy-Schwarz inequality:

$$\begin{aligned} \left| \langle u_{1}, \rho_{2,N_{2}} \rangle_{(*)} + \langle \rho_{1,N_{1}}, u_{2} \rangle_{(*)} - \langle \rho_{1,N_{1}}, \rho_{2,N_{2}} \rangle_{(*)} \right| \\ \leq \| u_{1} \|_{(*)} \| \rho_{2,N_{2}} \|_{(*)} + \| u_{2} \|_{(*)} \| \rho_{1,N_{1}} \|_{(*)} \\ + \| \rho_{1,N_{1}} \|_{(*)} \| \rho_{2,N_{2}} \|_{(*)}. \end{aligned}$$

$$(85)$$

From the conditions $\lim_{N\to\infty} \|\rho_{1,N}\|_{(*)} = 0$ and $\lim_{N\to\infty} \|\rho_{2,N}\|_{(*)} = 0$, the right-hand side of (85) converges to zeros as $N_1 \to \infty$ and $N_1 \to \infty$. Therefore, we obtain

$$\lim_{N_1 \to \infty} \lim_{N_2 \to \infty} \left(\langle u_1, \rho_{2,N_2} \rangle_{(*)} + \langle \rho_{1,N_1}, u_2 \rangle_{(*)} - \langle \rho_{1,N_1}, \rho_{2,N_2} \rangle_{(*)} \right) = 0.$$
(86)

Then, (36) is proved by the following equality.

$$\langle u_{1}, u_{2} \rangle_{(*)} - \sum_{\nu_{1}, \mu_{1}} \sum_{\nu_{2}, \mu_{2}} w_{\nu_{1}, \mu_{1}}^{\nu_{2}, \mu_{2}} \mathring{u}_{1, \nu_{1}, \mu_{1}}^{\text{int}}(\boldsymbol{r}_{c})^{*} \mathring{u}_{2, \nu_{2}, \mu_{2}}^{\text{int}}(\boldsymbol{r}_{c})$$

$$= \lim_{N_{1} \to \infty} \lim_{N_{2} \to \infty} \left(\langle u_{1}, \rho_{2, N_{2}} \rangle_{(*)} + \langle \rho_{1, N_{1}}, u_{2} \rangle_{(*)} - \langle \rho_{1, N_{1}}, \rho_{2, N_{2}} \rangle_{(*)} \right)$$

$$= 0.$$

$$(87)$$

Appendix B Interior Weights in Case of Interior Uniformly Weighted L_2 Norm

First, the inner product corresponding to the norm $\|\cdot\|_{int,uni}$ is given as

$$\langle u_1, u_2 \rangle_{\text{int,uni}} = \int_{\Omega_{\text{int}}} u_1(\boldsymbol{r})^* u_2(\boldsymbol{r}) \,\mathrm{d}\boldsymbol{r}.$$
 (88)

Therefore, the interior weights $w_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ are rewritten as

$$w_{\nu_{1},\mu_{1}}^{\nu_{2},\mu_{2}} = \int_{\Omega_{\text{int}}} \varphi_{\nu_{1},\mu_{1}} (\boldsymbol{r} - \boldsymbol{r}_{\text{c}})^{*} \varphi_{\nu_{2},\mu_{2}} (\boldsymbol{r} - \boldsymbol{r}_{\text{c}}) \,\mathrm{d}\boldsymbol{r}$$

$$= 4\pi \int_{0}^{R_{\text{int}}} j_{\nu_{1}} (kr)^{2} r^{2} \,\mathrm{d}r$$

$$\cdot \int_{\mathbb{S}^{2}} Y_{\nu_{1},\mu_{1}} (\hat{\boldsymbol{r}})^{*} Y_{\nu_{2},\mu_{2}} (\hat{\boldsymbol{r}}) \,\mathrm{d}\hat{\boldsymbol{r}}.$$
(89)

From the orthonormality of the spherical harmonics, one obtains

$$\int_{\mathbb{S}^2} Y_{\nu_1,\mu_1}(\hat{\boldsymbol{r}})^* Y_{\nu_2,\mu_2}(\hat{\boldsymbol{r}}) \,\mathrm{d}\hat{\boldsymbol{r}} = \delta_{\nu_1,\nu_2} \delta_{\mu_1,\mu_2}. \tag{90}$$

In addition, from the identity [35]:

$$\frac{\mathrm{d}}{\mathrm{d}r} \left[\frac{r^3}{2} \left(j_{\nu}(kr)^2 - j_{\nu-1}(kr) j_{\nu+1}(kr) \right) \right] = j_{\nu}(kr)^2 r^2, \quad (91)$$

one obtains

$$\int_{0}^{R_{\rm int}} j_{\nu}(kr)^{2} r^{2} dr$$

$$= \frac{R_{\rm int}^{3}}{2} \left(j_{\nu}(kR_{\rm int})^{2} - j_{\nu-1}(kR_{\rm int}) j_{\nu+1}(kR_{\rm int}) \right). \quad (92)$$

Equations (90) and (92) yield (46).

APPENDIX C Interior Weights in Case of Gaussian-Weighted L_2 Norm

The inner product corresponding to the norm $\|\cdot\|_{\mathrm{int,Gs}}$ is given as

$$\langle u_1, u_2 \rangle_{\text{int,Gs}} = \int_{\Omega_{\text{int}}} \exp\left(-\frac{\|\boldsymbol{r} - \boldsymbol{r}_c\|^2}{2\sigma^2}\right) u_1(\boldsymbol{r})^* u_2(\boldsymbol{r}) \,\mathrm{d}\boldsymbol{r}.$$
(93)

Therefore, the interior weights $w_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ are rewritten in a similar way to (89) as

$$w_{\nu_1,\mu_1}^{\nu_2,\mu_2} = \delta_{\nu_1,\nu_2} \delta_{\mu_1,\mu_2} 4\pi \int_0^{R_{\rm int}} \exp\left(-\frac{r^2}{2\sigma^2}\right) j_{\nu_1}(kr)^2 r^2 \,\mathrm{d}r.$$
(94)

Therefore, (49) is yielded by the following equation:

$$\frac{\mathrm{d}}{\mathrm{d}r} \left[\sum_{n=0}^{\infty} (-1)^n \frac{r^3}{2} \exp\left(-\frac{r^2}{4\sigma^2}\right) i_n\left(\frac{r^2}{4\sigma^2}\right) \\ \cdot \left(j_{\nu-n}(kr)j_{\nu+n}(kr) - j_{\nu-n-1}(kr)j_{\nu+n+1}(kr)\right) \right] \\ = \exp\left(-\frac{r^2}{2\sigma^2}\right) j_{\nu}(kr)^2 r^2.$$
(95)

Next, we prove (95). First, let $a_{n,\nu}(r)$ be defined as

$$a_{n,\nu}(r) = (-1)^n \frac{r^3}{2} \exp\left(-\frac{r^2}{4\sigma^2}\right) i_n\left(\frac{r^2}{4\sigma^2}\right) \cdot (j_{\nu-n}(kr)j_{\nu+n}(kr) - j_{\nu-n-1}(kr)j_{\nu+n+1}(kr)).$$
(96)

The derivative of $a_{n,\nu}(r)$ is written as

$$\begin{aligned} \frac{\mathrm{d}}{\mathrm{d}r} a_{n,\nu}(r) \\ &= (-1)^n r^2 \exp\left(-\frac{r^2}{4\sigma^2}\right) \frac{r^2}{4\sigma^2} \\ &\cdot \left\{ \left[\frac{n+1}{r^2/(4\sigma^2)} i_n\left(\frac{r^2}{4\sigma^2}\right) - i_n\left(\frac{r^2}{4\sigma^2}\right) + i'_n\left(\frac{r^2}{4\sigma^2}\right) \right] \\ &\quad \cdot j_{\nu-n}(kr) j_{\nu+n}(kr) \\ &\quad + \left[\frac{n}{r^2/(4\sigma^2)} i_n\left(\frac{r^2}{4\sigma^2}\right) + i_n\left(\frac{r^2}{4\sigma^2}\right) - i'_n\left(\frac{r^2}{4\sigma^2}\right) \right] \\ &\quad \cdot j_{\nu-n-1}(kr) j_{\nu+n+1}(kr) \right\} \\ &= (-1)^n r^2 \exp\left(-\frac{r^2}{4\sigma^2}\right) \frac{r^2}{4\sigma^2} \\ &\quad \cdot \left\{ \left[i_{n-1}\left(\frac{r^2}{4\sigma^2}\right) - i_n\left(\frac{r^2}{4\sigma^2}\right)\right] j_{\nu-n}(kr) j_{\nu+n}(kr) \\ &\quad + \left[i_n\left(\frac{r^2}{4\sigma^2}\right) - i_{n+1}\left(\frac{r^2}{4\sigma^2}\right)\right] j_{\nu-n-1}(kr) j_{\nu+n+1}(kr) \right\} \\ &= b_{n,\nu}(r) - b_{n+1,\nu}(r), \end{aligned}$$

where

$$b_{n,\nu}(r) = (-1)^n r^2 \exp\left(-\frac{r^2}{4\sigma^2}\right) \frac{r^2}{4\sigma^2}$$
$$\cdot \left[i_{n-1}\left(\frac{r^2}{4\sigma^2}\right) - i_n\left(\frac{r^2}{4\sigma^2}\right)\right] j_{\nu-n}(kr) j_{\nu+n}(kr).$$
(98)

From the inequality:

$$|j_{\nu}(x)| \le |i_{\nu}(x)| \le \left(\frac{x}{2}\right)^{\nu} \frac{\sqrt{\pi}}{2\Gamma\left(\nu + \frac{3}{2}\right)} \exp\left(\frac{x^2}{4}\right),$$
 (99)

one obtains the following relation in the sense of the compact convergence for $r \ge 0$:

$$\lim_{n \to \infty} b_{n,\nu}(r) = 0. \tag{100}$$

Equation (99) can be shown from the series representation of the Bessel functions [23], [33]. Therefore, one obtains

$$\frac{\mathrm{d}}{\mathrm{d}r} \sum_{n=0}^{\infty} a_{n,\nu}(r) = \sum_{n=0}^{\infty} \frac{\mathrm{d}}{\mathrm{d}r} a_{n,\nu}(r)$$
$$= b_{0,\nu}(r)$$
$$= \exp\left(-\frac{r^2}{2\sigma^2}\right) j_{\nu}(kr)^2 r^2.$$
(101)

This yields (95) and therefore (49).

Appendix D Exterior Weights in Case of Exterior Uniformly Weighted L_2 Norm

The inner product corresponding to the norm $\|\cdot\|_{\mathrm{ext,uni}}$ is given as

$$\langle u_1, u_2 \rangle_{\text{ext,uni}} = \int_{\Omega'_{\text{ext}}} u_1(\boldsymbol{r})^* u_2(\boldsymbol{r}) \,\mathrm{d}\boldsymbol{r}.$$
 (102)

Therefore, the exterior weights $v^{\nu_2,\mu_2}_{\nu_1,\mu_1}$ are rewritten as

$$v_{\nu_1,\mu_1}^{\nu_2,\mu_2} = \delta_{\nu_1,\nu_2} \delta_{\mu_1,\mu_2} 4\pi \int_{R_{\text{ext}}}^{R_{\text{ext}}'} |h_{\nu_1}(kr)|^2 r^2 \,\mathrm{d}r.$$
(103)

From the identity:

$$\frac{\mathrm{d}}{\mathrm{d}r} \left[\frac{r^3}{2} \left(h_{\nu}(kr)^* h_{\nu}(kr) - h_{\nu-1}(kr)^* h_{\nu+1}(kr) \right) \right]$$
$$= h_{\nu}(kr)^* h_{\nu}(kr) r^2, \tag{104}$$

which can be derived similarly to (91), one obtains

$$\int_{R_{\text{ext}}}^{R_{\text{ext}}} |h_{\nu}(kr)|^{2} r^{2} \, \mathrm{d}r$$

$$= \frac{R_{\text{ext}}^{\prime 3}}{2} \left(|h_{\nu}(kR_{\text{ext}})|^{2} - h_{\nu-1}(kR_{\text{ext}})^{*}h_{\nu+1}(kR_{\text{ext}}) \right)$$

$$- \frac{R_{\text{ext}}^{3}}{2} \left(|h_{\nu}(kR_{\text{ext}})|^{2} - h_{\nu-1}(kR_{\text{ext}})^{*}h_{\nu+1}(kR_{\text{ext}}) \right).$$
(105)

This yields (54).

APPENDIX E EXTERIOR WEIGHTS IN CASE OF EXTERIOR RADIATION-POWER NORM

The inner product corresponding to the norm $\|\cdot\|_{\mathrm{ext,rad}}$ is given as

$$\langle u_1, u_2 \rangle_{\text{ext,rad}}$$

$$=\frac{\mathrm{j}}{4\rho ck}\int_{S_{\mathrm{ext}}} \left[u_2(\boldsymbol{r})\frac{\partial}{\partial \hat{\boldsymbol{n}}(\boldsymbol{r})}u_1(\boldsymbol{r})^* - u_2(\boldsymbol{r})^*\frac{\partial}{\partial \hat{\boldsymbol{n}}(\boldsymbol{r})}u_1(\boldsymbol{r})\right]\mathrm{d}\boldsymbol{r}.$$
(106)

Therefore, the exterior weights $v_{\nu_1,\mu_1}^{\nu_2,\mu_2}$ are rewritten as

$$v_{\nu_{1},\mu_{1}}^{\nu_{2},\mu_{2}} = \delta_{\nu_{1},\nu_{2}} \delta_{\mu_{1},\mu_{2}} \frac{\pi j R_{\text{ext}}^{2}}{\rho c} \\ \cdot (h_{\nu_{1}}(kR_{\text{ext}})h_{\nu_{1}}'(kR_{\text{ext}})^{*} - h_{\nu_{1}}(kR_{\text{ext}})^{*}h_{\nu_{1}}'(kR_{\text{ext}})) \\ = \delta_{\nu_{1},\nu_{2}} \delta_{\mu_{1},\mu_{2}} \frac{2\pi}{\rho ck^{2}},$$
(107)

where the second line is derived using the Wronskian of the spherical Hankel functions [33]. This yields (57).

Next, we give the proof of (62) and (63). As shown in [24], the following two identities hold for $T_{\nu_1,\mu_1}^{\nu_2,\mu_2}(\cdot)$:

$$T^{\nu_2,\mu_2}_{\nu_1,\mu_1}(-\boldsymbol{r}) = T^{\nu_1,\mu_1}_{\nu_2,\mu_2}(\boldsymbol{r})^*,$$
(108)

$$T_{\nu_{1},\mu_{1}}^{\nu_{2},\mu_{2}}(\boldsymbol{r}+\boldsymbol{r}') = \sum_{\nu,\mu} T_{\nu_{1},\mu_{1}}^{\nu,\mu}(\boldsymbol{r}) T_{\nu,\mu}^{\nu_{2},\mu_{2}}(\boldsymbol{r}').$$
(109)

Therefore, one obtains

$$\begin{aligned} (\mathbf{A})_{l_{1},l_{2}} &= \frac{2\pi}{\rho c k^{2}} \sum_{\nu,\mu} \mathring{g}_{l_{1},\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{l_{1}})^{*} \mathring{g}_{l_{2},\nu,\mu}^{\text{ext}}(\boldsymbol{r}_{l_{2}}) \\ &= \frac{2\pi}{\rho c k^{2}} \sum_{\nu,\mu} \left(\sum_{\nu_{1},\mu_{1}}^{N_{l_{1}}} T_{\nu,\mu}^{\nu_{1},\mu_{1}}(\boldsymbol{r}_{c}-\boldsymbol{r}_{l_{1}})\beta_{l_{1},\nu_{1},\mu_{1}} \right)^{*} \\ &\cdot \left(\sum_{\nu_{2},\mu_{2}}^{N_{l_{2}}} T_{\nu,\mu}^{\nu_{2},\mu_{2}}(\boldsymbol{r}_{c}-\boldsymbol{r}_{l_{2}})\beta_{l_{2},\nu_{2},\mu_{2}} \right) \\ &= \frac{2\pi}{\rho c k^{2}} \sum_{\nu_{1},\mu_{1}}^{N_{l_{1}}} \sum_{\nu_{2},\mu_{2}}^{N_{l_{2}}} \beta_{l_{1},\nu_{1},\mu_{1}}^{*}\beta_{l_{2},\nu_{2},\mu_{2}} \\ &\sum_{\nu,\mu} T_{\nu,\mu}^{\nu_{1},\mu_{1}}(\boldsymbol{r}_{c}-\boldsymbol{r}_{l_{1}})^{*} T_{\nu,\mu}^{\nu_{2},\mu_{2}}(\boldsymbol{r}_{c}-\boldsymbol{r}_{l_{2}}) \\ &= \frac{2\pi}{\rho c k^{2}} \sum_{\nu_{1},\mu_{1}}^{N_{l_{1}}} \sum_{\nu_{2},\mu_{2}}^{N_{l_{2}}} \beta_{l_{1},\nu_{1},\mu_{1}}^{*}\beta_{l_{2},\nu_{2},\mu_{2}} T_{\nu,\mu}^{\nu_{2},\mu_{2}}(\boldsymbol{r}_{l_{2}}-\boldsymbol{r}_{l_{2}}). \end{aligned}$$
(110)

This yields (62). Similarly, (63) can be obtained.

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