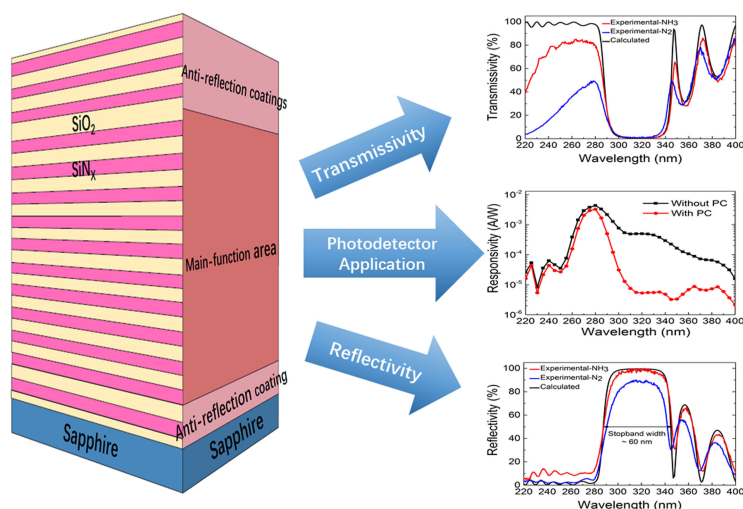


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# A High-Performance SiO<sub>2</sub>/SiN<sub>x</sub> 1-D Photonic Crystal UV Filter Used for Solar-Blind Photodetectors

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**Abstract:** To improve the solar-blind/visible-blind photocurrent response rejection ratio of solar-blind photodetectors, we designed and fabricated a high-performance SiO<sub>2</sub>/SiN<sub>x</sub> 1-D photonic crystal (PC) ultraviolet (UV) filter on (0 0 0 1) double-polished sapphire substrate. When depositing SiN<sub>x</sub>, we found that employing NH<sub>3</sub> as nitrogen precursor instead of N<sub>2</sub> can simultaneously improve the peak reflectivity of filter stopband in the designed visible-blind region and the transmissivity in the solar-blind region. Research shows that it is associated with the H atom concentration in SiN<sub>x</sub> and the generation of Si-O-N transition layer at SiO<sub>2</sub>/SiN<sub>x</sub> interface. Finally, we obtained a high-performance SiO<sub>2</sub>/SiN<sub>x</sub> PC UV filter with a stopband reflectivity over 90% from 285 to 345 nm and a transmissivity over 80% in the solar-blind region. The UV filter is also demonstrated to have a good effect in improving the solar-blind/visible-blind photocurrent response rejection ratio of a back-illuminated AlGaIn photodetector by depositing it to the back of the detector.

**Index Terms:** Photonic crystal (PC), UV Filter, SiO<sub>2</sub>/SiN<sub>x</sub>, photodetectors.

## 1. Introduction

Solid state solar-blind ultraviolet (UV) photodetectors are attracting more and more attention because of their potential applications in civil and military fields such as engine control, UV deep space exploration, flame sensors, missile warning and secure space-to-space communication [1]–[5]. Out-of-band light response of solar-blind region is still a key problem that influences the detecting sensitivity of devices to weak solar-blind UV light (<290 nm) signal, although some materials, such as AlGaIn and Ga<sub>2</sub>O<sub>3</sub> wide bandgap semiconductors widely concerned in recent, are of characteristic of intrinsic solar-blind. Detectors made out of these materials would be sensitive only to ultraviolet light that does not penetrate the earth’s atmosphere. However, defects and impurities that can result in photocurrent response from out-band of solar-blind in the wide bandgap

TABLE 1  
Deposition Parameters of SiO<sub>2</sub> and SiN<sub>x</sub> Layers with NH<sub>3</sub> and N<sub>2</sub> as Nitrogen Precursor, Respectively, by PECVD

Deposition material	SiO <sub>2</sub>	SiN <sub>x</sub> (N <sub>2</sub> )	SiN <sub>x</sub> (NH <sub>3</sub> )
Growth rate/nm min <sup>-1</sup>	21.41	6.52	8.95
Reaction temperature/°C	350	350	350
RF power/W	10	15	50
Reaction pressure/mTorr	300	600	500
N <sub>2</sub> O gas flow/cm <sup>3</sup> min <sup>-1</sup>	400	/	/
N <sub>2</sub> gas flow/cm <sup>3</sup> min <sup>-1</sup>	/	900	400
NH <sub>3</sub> gas flow/cm <sup>3</sup> min <sup>-1</sup>	/	900	20
SiH <sub>4</sub> /N <sub>2</sub> gas flow/cm <sup>3</sup> min <sup>-1</sup>	100	25	25

compounds are unavoidable. So the UV filters with high reflectivity in visible-blind region and high transmissivity in solar-blind region are still needed for developing the solar-blind photodetectors.

The 1-D photonic crystal (PC) filter based on SiO<sub>2</sub>/SiN<sub>x</sub> films can achieve very high reflectivity and wide stopband in theory because of relative large contrast of refractive index between SiO<sub>2</sub> and SiN<sub>x</sub> [6]–[8]. Meanwhile, periodic SiO<sub>2</sub>/SiN<sub>x</sub> films can be deposited directly on the back of back-illuminated photodetectors and hence this SiO<sub>2</sub>/SiN<sub>x</sub> based filter can be integrated easily to the photodetector. However, it is very challenging to grow high quality SiO<sub>2</sub>/SiN<sub>x</sub> based 1-D PC filter with high reflectivity in the visible-blind region and transmissivity in the solar-blind region simultaneously, because solar-blind wave band is very close to their optical bandgaps of SiO<sub>2</sub> and SiN<sub>x</sub>, and meanwhile reflectivity in the visible-blind region and transmissivity in the solar-blind region are two contradictory parameters that one requires more periodic layers to increase reflection to visible-blind light and the other requires less layers to decrease absorption to solar-blind light.

In this work, we designed and fabricated a 1-D PC UV filter to improve the solar-blind/visible-blind photocurrent response rejection ratio of solar-blind photodetectors, and the filter consists of periodic SiO<sub>2</sub>/SiN<sub>x</sub> (L/2 H L/2)<sub>m</sub> units sandwiched between two anti-reflection coatings. The structure and growth parameters of the UV filter including gas sources and flow rate were optimized by means of the analysis of optical properties, microstructure, and depth profiles of atom concentration as well as theoretical calculation. Finally, we obtained a high performance SiO<sub>2</sub>/SiN<sub>x</sub> 1-D PC UV filter with a stopband reflectivity over 90% from 285 nm to 345 nm and a transmissivity over 80% in the solar-blind region. In addition, the actual effect in improving solar-blind/visible-blind photocurrent response rejection ratio of the UV filter was demonstrated by integrating the filter to the back of an AlGaIn solar-blind photodetector.

## 2. Experimental

All SiO<sub>2</sub>/SiN<sub>x</sub> 1-D PC UV filter samples were fabricated by plasmas-enhanced chemical vapor deposition (OXFORD Plasma 80 Plus PECVD) on (0001) double-polished sapphire substrates. The silicon and oxygen precursors were SiH<sub>4</sub> (5%) and N<sub>2</sub>O, respectively. For comparison, we choose N<sub>2</sub> [9], [10] and NH<sub>3</sub> [11], [12] as nitrogen precursor, respectively. Initially, the single SiN<sub>x</sub> and SiO<sub>2</sub> layers were deposited respectively at the reactor to calibrate the growth rate and optical constants about refractive index (n) and extinction coefficient (k). The detailed experimental parameters are listed in Table 1. The growth rate was determined by measuring the thickness of the films using a step profiler.

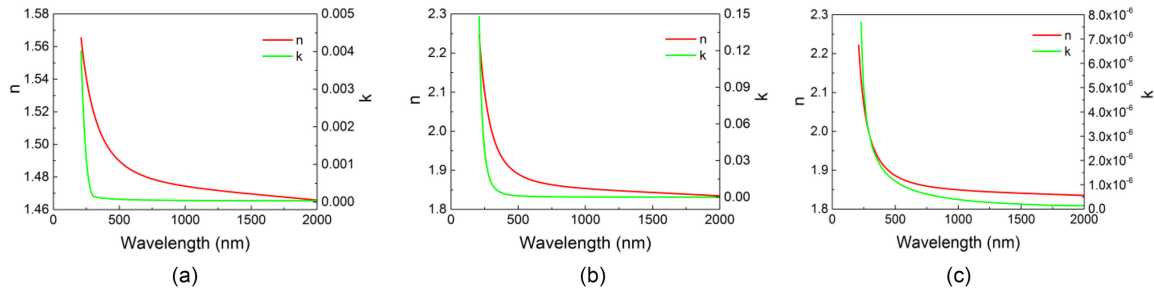


Fig. 1. Refractive index ( $n$ ) and extinction coefficient ( $k$ ) of SiO<sub>2</sub> (a) and SiN<sub>x</sub> grown with N<sub>2</sub> (b) and NH<sub>3</sub> (c) as nitrogen precursor measured by spectroscopic ellipsometer.

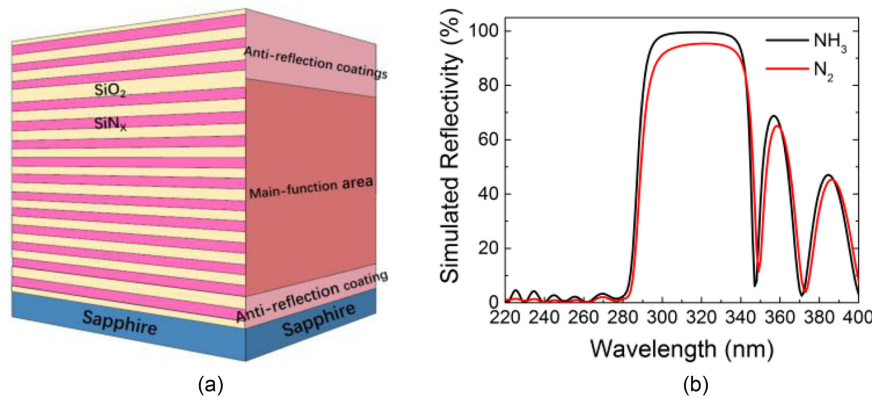


Fig. 2. Schematic diagram (a) and the calculated reflectivity spectra (b) of designed UV filters considering N<sub>2</sub> and NH<sub>3</sub> as nitrogen precursor, respectively.

In addition, the structures of UV filter were deposited according to the calibrated conditions. The reflectivity and transmissivity spectra were measured by UV-visible spectrometer (SHIMADZU UV-3600). The surface morphologies of samples were observed by atomic force microscope (NT-MDT Modular AFM). The cross-sectional images of samples were characterized by transmission electron microscope (FEI Tecnai G2 F20 S-TWIN TEM). The concentration depth profiles of elements in the samples were measured by secondary ion mass spectroscopy (PHI Adept 1010 SIMS).

### 3. Structure Design and Optimization

A center wavelength of 310 nm and a stopband from 290 to 350 nm are expected for UV filters used for solar-blind photodetectors. According to this requirement, we optimized the structure of the PC UV filter by using the transfer matrix method based on the experimental optical constants of SiO<sub>2</sub> and SiN<sub>x</sub>. The refractive index and extinction coefficient were measured by the EOPTICS ME-L broadband spectroscopic ellipsometer, as shown in Fig. 1(a), (b), (c). It is noteworthy that the SiN<sub>x</sub> layer grown with NH<sub>3</sub> as nitrogen precursor has a significant reduction in extinction coefficient by comparing Fig. 1(b) and Fig. 1(c). The optimized structure of 1-D PC UV filter consists of the periodic main-function units (L/2 H L/2)<sub>m</sub> and two anti-reflection coatings, as illustrated in Fig. 2(a). L and H represent the thickness of SiO<sub>2</sub> and SiN<sub>x</sub>, respectively. The main-function units consist of 10-period SiO<sub>2</sub>/SiN<sub>x</sub>/SiO<sub>2</sub> layers, and the thickness of three layers is 25.5/39.3/25.5 nm calculated by

$$d = \lambda_0/4n$$

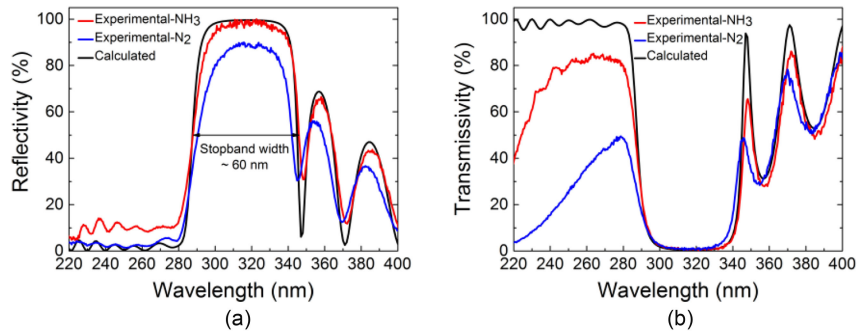


Fig. 3. The measured reflectivity (a) and transmissivity (b) spectra of the UV filters grown with NH<sub>3</sub> and N<sub>2</sub> as nitrogen precursor, respectively. The calculated reflectivity and transmissivity spectra are also presented as reference.

Two anti-reflection coatings are similar to the main-function units besides a slight change in the thickness of each layer. The top anti-reflection structure is  $1.04 \cdot (L/2 \ H \ L/2)_3$  and the bottom one is  $1.35 \cdot (L/2 \ H \ L/2)$ . The reflectivity spectra calculated by transfer matrix method and simulated by TFCalc system are shown in Fig. 2(b). The calculated spectra according to the measured optical constants show that the UV filter fabricated with NH<sub>3</sub> as nitrogen precursor has a higher reflectivity of approaching 100% compared to the UV filter fabricated with N<sub>2</sub> as nitrogen precursor. The calculated stopband covers from about 285 nm to 340 nm for both the UV filters.

#### 4. Results and Discussion

Fig. 3 shows the actual reflectivity and transmissivity spectra of the SiO<sub>2</sub>/SiN<sub>x</sub> PC UV filters fabricated using different nitrogen precursors, measured by UV-visible spectrometer. The filter using NH<sub>3</sub> as nitrogen precursor has a higher maximum stopband peak reflectivity of 99% and a wider stopband width of 60 nm from 285 nm to 345 nm, although the reflectivity in the solar-blind region is slightly higher when comparing with that using N<sub>2</sub> as nitrogen precursor. It can be seen that the experimental reflectivity spectrum of the optimized UV filter is in a good coincidence with the calculated reflectivity spectrum. Furtherly, the transmissivity of the filter using NH<sub>3</sub> as nitrogen precursor in the solar-blind region is twice stronger than that of the filter using N<sub>2</sub> as nitrogen precursor, as shown in Fig. 3(b), indicating that there exists very serious light absorption in the solar-blind region for the filter using N<sub>2</sub> as nitrogen precursor.

To explain the causes that result in the large discrepancies on reflectivity and transmissivity of the SiO<sub>2</sub>/SiN<sub>x</sub> PC UV filters fabricated using different nitrogen precursors, we analyzed the element depth profiles of the two kinds of UV filters. Fig. 4(a) and (b) show depth profiles of ion concentration for Si, O, N, and H in the SiO<sub>2</sub>/SiN<sub>x</sub> PC UV filters grown with N<sub>2</sub> and NH<sub>3</sub> as nitrogen precursor, respectively, measured by SIMS. On the whole, depth profile curves present periodic composition fluctuations in accordance with the designed SiO<sub>2</sub>/SiN<sub>x</sub> bilayer structure, and the discrepancies on the ion concentrations of Si, N, and H in the SiN<sub>x</sub> layers of the two UV filters can be observed obviously. The SiN<sub>x</sub> layer grown with N<sub>2</sub> as nitrogen precursor exhibits the feature of N-rich SiN<sub>x</sub> with a N/Si ratio of about 5:3, and has a high H atom concentration of about 16%. In contrast, the SiN<sub>x</sub> layer grown with NH<sub>3</sub> as nitrogen precursor exhibits the feature of Si-rich SiN<sub>x</sub> with a N/Si ratio of almost 1:1, and has a low H atom concentration of about 2%. To investigate the effect of intrinsic absorption from the N-rich SiN<sub>x</sub> and Si-rich SiN<sub>x</sub> films on the reflectivity and transmissivity spectra of the SiO<sub>2</sub>/SiN<sub>x</sub> PC UV filters, we extracted optical bandgap of the SiN<sub>x</sub> films by characterizing the transmissivity spectra with the plots [13]–[15], as shown in Fig. 5. The extracted optical bandgaps are 5.14 eV and 5.31 eV for the N-rich SiN<sub>x</sub> and Si-rich SiN<sub>x</sub> films, respectively, corresponding to the intrinsic absorption cut-off edge of 241 nm to 234 nm. This means that the intrinsic absorption of SiN<sub>x</sub> films has little effect on the reflectivity and transmissivity spectra of the PC UV filters in

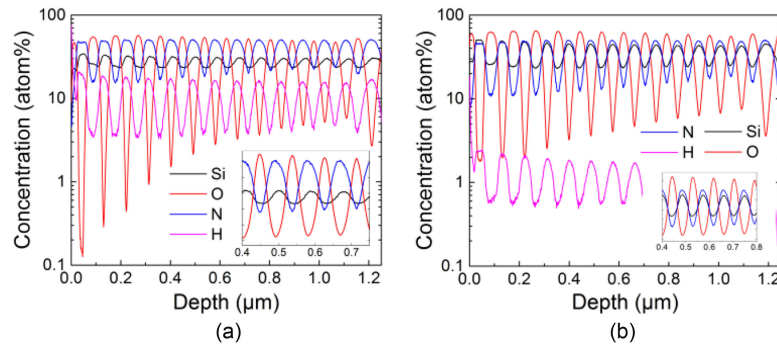


Fig. 4. Depth profile of ion concentration for Si, O, N, and H in the UV filter structure grown with N<sub>2</sub> (a) and NH<sub>3</sub> (b) as nitrogen precursor, respectively, measured by SIMS.

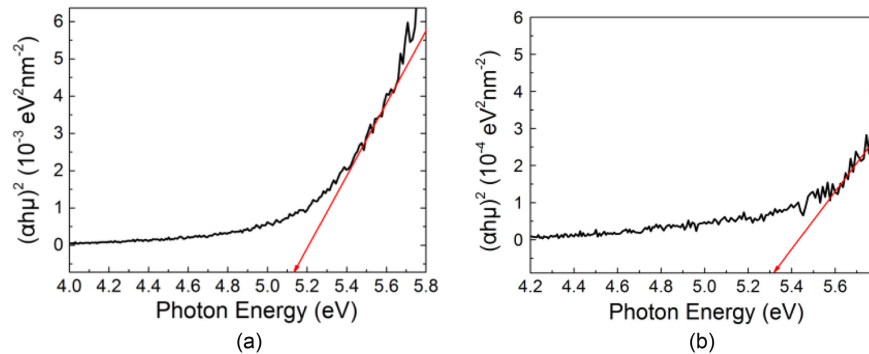


Fig. 5. The optical bandgap characterized from the transmissivity spectra of SiN<sub>x</sub> grown with N<sub>2</sub> (a) and NH<sub>3</sub> (b) as nitrogen precursor, respectively.

the wavelength range of 240–345 nm. So, we think that the high H atom concentration in the N-rich SiN<sub>x</sub> film is responsible for its decreasing reflectivity in the stopband and transmissivity in the solar-blind region. This can be explained by the absorption of Si-H bond to the photons with energy above 3.52 eV [16], [17], which is also in good agreement with the large discrepancy in extinction coefficient between the SiN<sub>x</sub> films grown with NH<sub>3</sub> and N<sub>2</sub> as nitrogen precursor, respectively. However, the absorption of Si-H bond cannot explain the larger transmissivity discrepancy in the solar-blind region between two PC UV filters. This large transmissivity discrepancy can be mainly attributed to the formation of a thin Si-O-N transition layer between the SiO<sub>2</sub> and SiN<sub>x</sub> layers in the PC UV filter with N<sub>2</sub> as nitrogen precursor. The Si-O-N transition layer with graded composition can be observed obviously from the inset of Fig. 4(a), and this transition layer was also often observed in other references [18], [19]. The Si-O-N film has an extensive range of optical bandgap from 2.9 to 6.2 eV according to different references [20], [21], and hence exists significant absorption in the solar-blind region.

$$(\alpha h\mu)^2 - h\mu$$

Fig. 6(a) and Fig. 6(b) show the surface morphology measured by AFM and cross-sectional TEM image of the SiO<sub>2</sub>/SiN<sub>x</sub> PC UV filter using NH<sub>3</sub> as nitrogen precursor. A smooth surface with altitude range less than 16 nm is observed, and the root mean square roughness is only 1.85 nm which is much lower than the structure with the similar thickness reported in recent literature [22]. In the TEM image, we can observe a 29-layer periodic structure composed of two alternating materials where the white layer represents SiO<sub>2</sub> and the black layer represents SiN<sub>x</sub>. The SiO<sub>2</sub>/SiN<sub>x</sub> interfaces near

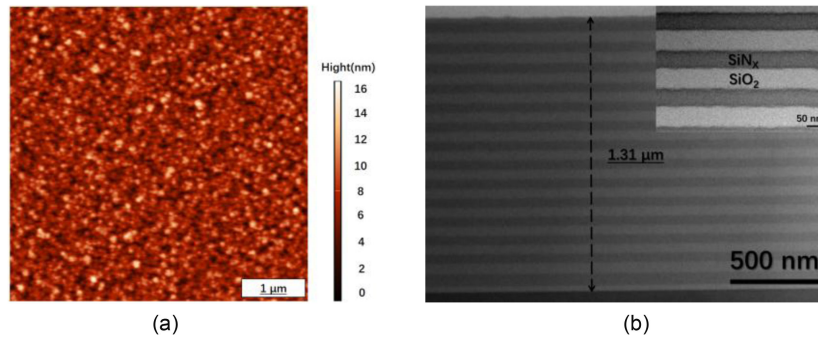


Fig. 6. Surface morphology (a) and cross-sectional image (b) of UV filter grown with NH<sub>3</sub> as nitrogen precursor revealed by AFM and TEM, respectively.

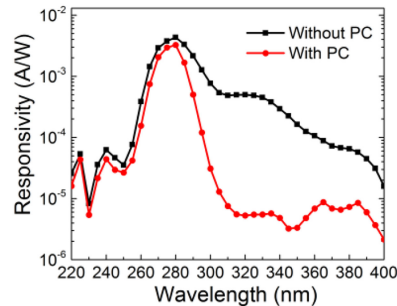


Fig. 7. Spectral responsivities of the p-Al<sub>0.2</sub>Ga<sub>0.8</sub>N/i-Al<sub>0.4</sub>Ga<sub>0.6</sub>N/n-Al<sub>0.4</sub>Ga<sub>0.6</sub>N solar-blind photodetector under back illumination in both cases before and after depositing PC UV filter.

sapphire substrate are very clear and pretty flat, and become slightly unsmooth with the increase of the number of layers of filter as shown in the inset. The whole thickness of the UV filter measured is 1.31  $\mu\text{m}$ , and the mean thicknesses of SiO<sub>2</sub> and SiN<sub>x</sub> in the main function area are 51.3 nm and 39.1 nm, respectively. The measured thickness from the TEM image agrees well with the designed one with a periodic 51.0 nm/39.3 nm SiO<sub>2</sub>/SiN<sub>x</sub> structure.

To verify the actual effect of the SiO<sub>2</sub>/SiN<sub>x</sub> PC UV filter, we deposited it to the back of a traditional p-Al<sub>0.2</sub>Ga<sub>0.8</sub>N/i-Al<sub>0.4</sub>Ga<sub>0.6</sub>N/n-Al<sub>0.4</sub>Ga<sub>0.6</sub>N solar-blind photodetector. Fig. 7 shows the photocurrent responsivity of the solar-blind photodetector under back illumination in both cases before and after depositing PC UV filter by using a calibrated monochromator and a Xe arc lamp as light source. The power of the monochromatic light was tested by a calibrated silicon. It can be seen that the spectral responsivity of the solar-blind photodetector integrated with PC UV filter presents a sharper cutoff wavelength at 285 nm and the response peak at 330 nm resulted from the photo-generated current of the p-Al<sub>0.2</sub>Ga<sub>0.8</sub>N layer disappears in comparison with the as-fabricated solar-blind photodetector. Although the responsivity in the solar-blind region decreases slightly, the photocurrent response from out of solar-blind region is significantly suppressed by almost two orders of magnitude. We attribute the suppression of photocurrent response from out of solar-blind region to the high reflectivity of the 1-D PC UV filter in the visible-blind region.

## 5. Conclusion

In summary, a high performance SiO<sub>2</sub>/SiN<sub>x</sub> 1-D photonic crystal UV filter with stopband reflectivity over 90% from 285 nm to 345 nm and a transmissivity over 80% in the solar-blind region has been successfully fabricated by optimizing the deposition parameters and structure. This can be attributed to the significant reduction in extinction coefficient and the suppression of the generation

of Si-O-N transition layer at SiO<sub>2</sub>/SiN<sub>x</sub> interface when using NH<sub>3</sub> as nitrogen precursor instead of N<sub>2</sub>. When integrating the UV filter to the back of a AlGaIn solar-blind photodetector, the photocurrent response of the detector from out of solar-blind region can be significantly suppressed by almost two orders of magnitude.

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