

# The Kadomtsev Pinch Revisited for Sheared-Flow-Stabilized Z-Pinch Modeling

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**Abstract**—The Kadomtsev pinch, namely, the Z-pinch profile marginally stable to interchange modes, is revisited in light of observations from axisymmetric magnetohydrodynamics (MHD) modeling of the fusion Z-pinch experiment (FuZE) sheared-flow-stabilized Z-pinch experiment. We show that Kadomtsev’s stability criterion, cleanly derived by the minimum energy principle but of opaque physical significance, has an intuitive interpretation in the specific entropy analogous to the Schwarzschild–Ledoux criterion for convective stability of adiabatic pressure distributions in the fields of astrophysics, meteorology, and oceanography. By analogy, the Kadomtsev profile may be described as magnetoadiabatic in the sense that plasma pressure is polytropically related to area-averaged current density from the ideal MHD stability condition on the specific entropy. Furthermore, the nonideal stability condition of the entropy modes is shown to relate the specific entropy gradient to the ideal interchange stability function. Hence, the combined activity of the ideal interchange and nonideal entropy modes drives both the specific entropy and specific magnetic flux gradients to zero in the marginally stable state. The physical properties of Kadomtsev’s pinch are reviewed in detail and following from this the localization of pinch confinement, i.e., pinch size and inductance, is quantified by the ratio of extensive magnetic and thermal energies. In addition, results and analysis of axisymmetric MHD modeling of the FuZE Z-pinch experiment are presented where pinch structure is found to consist of a near-marginal flowing core surrounded by a supermagnetoadiabatic low-beta sheared flow.

**Index Terms**—Accelerators, current distribution, entropy, fusion reactors, magnetohydrodynamics (MHD), plasma pinch, plasma stability.

## I. INTRODUCTION

THE self-field plasma confinement equilibrium, popularly known as the Z-pinch, magnetically confines plasma

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pressure without external magnetic coils. Studies during the early days of plasma physics identified magnetohydrodynamic (MHD) instabilities in pinches of axisymmetric and kink types, and the axisymmetric modes later fell under the broader category of interchange-type modes, including, for example, the tokamak ballooning modes. In fact, Kadomtsev’s seminal article “Hydrodynamic stability of a plasma” [1] demonstrated the existence of a Z-pinch equilibrium marginally stable to the interchange modes using the method of virtual displacement and the energy principle. For this reason, this equilibrium will be referred to here as “Kadomtsev’s pinch.” Kadomtsev’s article noted that the interchange modes are analogous to buoyancy-driven modes in stratified fluids (e.g., Rayleigh–Taylor type) and expressed the stability condition in terms of the logarithmic pressure gradient as a rational function of plasma  $\beta$  and adiabatic index  $\gamma$ . On the other hand, the same analysis concluded that the kink modes of azimuthal mode number  $m = 1$  were unstable for these interchange-stabilized distributions (or any other static distribution, for that matter).

In the years since Kadomtsev’s original work, the community’s understanding of stratified buoyant fluids and magnetohydrodynamic interchange modes has considerably deepened. This deepening has centered around two somewhat complex elements, namely, the role of entropy gradients and sheared flows. In revisiting Kadomtsev’s pinch, this work is intended to fill out the analogy between stable stratification and interchange stability, to clarify the role of specific entropy in defining the marginally stable state, and to shed light on the effect of sheared flow on interchange stability with evidence from axisymmetric resistive MHD simulations of zap energy’s FuZE device. The intention is to reinterpret Kadomtsev’s stability criterion from the somewhat nebulous, “that the pressure does not fall off too rapidly with radius,” expressed by the logarithmic pressure gradient as a function of local plasma  $\beta$  with an opaque physical significance, to an intuitive formulation in terms of the specific entropy and specific magnetic flux gradients.

This article is organized as follows. Section II reviews buoyant convection in fluid dynamics, the influence of sheared flows on interchange stability, and the Kadomtsev profile in the literature. A theoretical exposition on Kadomtsev’s pinch follows in Section III, which first reviews the closely related phenomena of the adiabatic atmosphere and isentropic flows. Kadomtsev’s pinch is studied in a physically motivated approach where the stability condition is expressed as a zero gradient of what some authors have termed magnetic

entropy [2]. The condition admits a first integral of the force balance which rearranges to Kadomtsev's implicit pressure profile, noted as analogous to meteorology's adiabatic pressure profile. Analytic properties are studied for the adiabatic indices  $\gamma = 2$ ,  $\gamma = 5/3$ , and  $\gamma \rightarrow 1$ . The analysis is applied to the unity average beta concept in Section III-G and used to quantify pinch radius as a function of the ratio of total magnetic-to-thermal energies. Finally, Section IV considers idealized theory for the flow velocity of a Z-pinch downstream from a coaxial accelerator and presents 2-D axisymmetric simulations of the FuZE experiment demonstrating self-organized flow pinch states where Kadomtsev's pinch profile plays a key role.

## II. INTERCHANGE MODES IN THE LITERATURE

The hydrodynamic stability of stratified fluids is an important topic in all the fluid dynamic disciplines with convective motions, including terrestrial and astrophysical meteorology [3], oceanography [4], stellar astrophysics [5], [6], and plasma physics where convective instabilities arise as a category of ideal MHD instability known as interchange modes [7]. In the original summary of ideal MHD stability analysis through the energy method, Kadomtsev [1] briefly observed that the physics of the interchange mode were analogous to that of general buoyant stratification. In Section III, this analogy is recapitulated and the detailed mathematics worked out.

On the other hand, the situation can first be summarized as follows. Like many fluid dynamical topics, convective instability is known as a complex phenomenon with a variety of control factors such as chemical composition [8], sheared flow [9], [10], radiation [11], magnetization [12], external heating (Rayleigh-Bénard), discontinuous profiles (Rayleigh-Taylor), and impulsive forcing by shock waves (Richtmyer-Meshkov). The situation with flows is far from simple, since while the interchange is stabilized by sheared flow, the second derivative of transverse velocity is destabilizing (Kelvin-Helmholtz), and the question of stability under shear is rather complex due to the prevalence of nonmodal solutions in the initial-value problem [13], the related topic of stabilization of pseudomodes in the pseudospectrum [14], and 3-D perturbations [15]. Sheared flows, however, are generally understood to be stabilizing and to allow the persistence of superadiabatic profiles across the fluid dynamic disciplines [16]. Section IV presents some simulation results and commentary on the role of sheared flows on the Kadomtsev Z-pinch profile.

Regardless of the various complexities discussed above, fortunately the stability of a continuously stratified medium subject only to pressure and buoyancy is simple. The marginally stable profile in atmospheric meteorology and astrophysics, known as the adiabatic atmosphere, is characterized by a vanishing gradient of the specific entropy  $s$  such that  $\nabla s \cdot \nabla p = 0$ , a result known as Schwarzschild's criterion [17]. This condition is widely used in astrophysics [18], meteorology [19], and plasma physics [20], [21]. The fundamental reason for the specific entropy to characterize the marginal state is that the displacement frequency due to buoyancy,

known as the Brunt-Väisälä frequency, is a function only of the invariants of motion along the displacement [22]. As the conserved quantity in an adiabatic displacement is the specific entropy, the local buoyancy frequency is a direct measure of the specific entropy gradient. In Section III-B, we will see how Kadomtsev's pressure profile is precisely the analogous "adiabatic" profile for the interchange-stable pinch, which is characterized by the invariance of a quantity which Pastukhov [2] and Kesner [7] refer to as magnetic entropy. It seems to the authors of this article that this entropy terminology does not correspond with that of Minardi [23], which is properly developed from an information-theoretic perspective on the distribution of electric current. Therefore this article avoids the terminology of magnetic entropy for the defining invariant of Kadomtsev's pinch, which is instead referred to as the quantity  $s_z$ .

The Kadomtsev-stable profiles make a frequent appearance in the Z-pinch literature. Kadomtsev profiles are observed as attractors in 2-D MHD simulations [24] and are used as initial conditions in simulations of sheared-flow stabilization [25], [26]. Although the Kadomtsev-stable profile is strictly unstable to the  $m = 1$  kink mode, under certain conditions it is observed experimentally. Chiefly, these are the hardcore Z-pinch and levitated dipole experiments which fix the magnetic axis with a conductor [27]. Furthermore, Kadomtsev's pinch has been invoked to explain data in compressional Z-pinch experiments [28], [29] and supporting theory [30]. Finally, it is worth noting that also in astrophysics, meteorology, and oceanography, adiabatic profiles are often invoked alongside other organizing principles, and sometimes do not explain observations at all, due to the myriad of complications discussed above. For this reason, the understanding of pressure, density, and temperature profiles in these disciplines remains an active research topic, as it certainly does in the field of magnetic plasma confinement. On the topic of self-organization in magnetic confinement with flows, the reader is pointed to work on double Beltrami fields and multiple-region relaxed magnetohydrodynamics [31], [32], [33], [34].

## III. PROPERTIES OF KADOMTSEV'S PINCH

### A. Analogy to Stable Stratification in the Adiabatic Atmosphere

Kadomtsev showed that the marginally stable pinch profile was implicitly a function of the local value of  $\beta = p/p_B$  with  $p$  the thermal pressure and  $p_B$  the magnetic pressure

$$\frac{p_0}{p} = \left(1 + \frac{\gamma - 1}{\gamma} \frac{2}{\beta}\right)^{\gamma/(\gamma-1)}. \quad (1)$$

Note that  $\beta$  is often defined in some integral manner, as discussed further in Section III-G, yet this quantity is the local plasma  $\beta$ . Now, to make the analogy to stable stratification in the atmosphere clear, consider the static equilibrium of a stratified gas of mass density  $\rho$  in a gravitational potential  $\Phi_g$

$$-\nabla p - \rho \nabla \Phi_g = 0. \quad (2)$$

Recall that the specific entropy  $s$  of a perfect gas is given by  $s = c_v \ln(p/\rho^\gamma)$  with  $c_v$  the specific heat at constant volume and  $\gamma$  the adiabatic index. Now noting that

$(p/p_0) = (\rho/\rho_0)^\gamma e^{s/c_v}$  with  $(p_0, \rho_0)$  reference constants making the logarithmic argument dimensionless, the pressure gradient is

$$\nabla p = c_s^2 \nabla \rho + p \nabla s / c_v \quad (3)$$

with  $c_s^2 = \gamma p / \rho$  the local speed of sound squared. Substituting into (2) gives another form of equilibrium

$$c_s^2 \nabla \rho = -\rho \nabla \Phi_g - (\gamma - 1) \rho T \nabla s. \quad (4)$$

Using the Schwarzschild condition  $\nabla s = 0$  means that the marginal state satisfies the equation

$$c_s^2 \frac{d\rho}{dz} = -\rho \frac{d\Phi_g}{dz}. \quad (5)$$

The solution to (5) is called the adiabatic atmosphere. Note that in the isentropic solution the sound speed  $c_s^2 = c_s^2(r)$  is variable unless one takes the isothermal limit of  $\gamma \rightarrow 1$ . Expressed solely as a function of density for a particular adiabat, the sound speed is given by

$$c_s^2 = c_{s0}^2 \left( \frac{\rho}{\rho_0} \right)^{\gamma-1} \quad (6)$$

where  $c_{s0}^2 = \gamma R T_0$ . We note that specifying  $\rho_0$  and  $T_0$  is sufficient to specify the constant  $e^{s_0/c_v}$  defining the adiabat of the atmosphere, since any constant specific entropy  $s_0$  satisfies the stability constraint. Defining  $\tilde{\rho} = \rho/\rho_0$ , we obtain the separable differential equation

$$\tilde{\rho}^{\gamma-2} \frac{d\tilde{\rho}}{dz} = -c_{s0}^{-2} \frac{d\Phi_g}{dz} \quad (7)$$

which with the initial condition  $\rho(0) = \rho_0$  integrates into

$$\frac{\rho}{\rho_0} = \left( 1 - \frac{\gamma-1}{\gamma} \frac{m\Phi_g}{k_B T_0} \right)^{1/(\gamma-1)} \quad (8)$$

with the corresponding pressure solution  $p = p_0(\rho/\rho_0)^\gamma$ . It is clear that Kadomtsev's pinch profile given by (1) is closely related to this pressure profile. Considering the isothermal limit  $\gamma \rightarrow 1$  provides us with the physical meaning of this solution

$$\rho = \rho_0 \exp\left(-\frac{m\Phi_g}{k_B T_0}\right) \quad (9)$$

as a Boltzmann distribution. For realistic values of  $\gamma$ , the isentropic solution is not isothermal, yet the marginal state is still nearly Boltzmann-distributed, albeit in the distribution's less familiar limiting form of  $e^x = \lim_{n \rightarrow \infty} (1 + x/n)^n$ . Therefore, in this work, we refer to the form of (8) as the frozen Boltzmann distribution. In the physical and mathematical literature, distributions in the form of (8) are referred to as  $\kappa$ - or  $q$ -Gaussian distributions and maximize a Tsallis-type entropy functional for the number density distribution [35].

Finally, observe that (8), the adiabatic atmosphere, is a first integral of the momentum equation, and thus expresses constancy of specific enthalpy

$$c_p T + \Phi_g = c_p T_0 \quad (10)$$

in the same way as the pressure–Mach number relationship for compressible isentropic ( $\nabla s = 0$ ) flow

$$\frac{p_0}{p} = \left( 1 + \frac{\gamma-1}{\gamma} \frac{mv^2/2}{k_B T_0} \right)^{\gamma/(\gamma-1)} \quad (11)$$

is a rearranged form of a constant specific enthalpy

$$c_p T + \frac{1}{2} v^2 = c_p T_0 \quad (12)$$

which similarly limits as  $\gamma \rightarrow 1$  to a Boltzmann distribution in the kinetic energy. In these particular cases, the adiabatic equilibrium as a frozen Boltzmann distribution can be understood to consist of both constant energy and entropy per particle, completely determining the distribution provided that there are only two forms of energy. Yet the ideal MHD picture is more complex as there are three forms of energy.

### B. Pinch Marginal Stability in Thermodynamic Variables

In this section, the marginal stability condition of the cylindrical plasma pinch to axisymmetric interchange modes is cast as a condition on the specific entropy and the specific magnetic flux gradients. This reformulation results in a particularly simple form of the stability criterion and the marginally stable state, in particular in light of the activity of the entropy modes discussed in Section III-C. The classical method to determine magnetohydrodynamic stability is accomplished analytically from the minimum energy principle provided that the force operator of a virtual displacement is self-adjoint [36]. According to this intuition, the reader is reminded of the equivalence in thermodynamics between the principle of minimum energy and the principle of maximum entropy [37]. Now, it is well-known that given such a virtual displacement of the static MHD plasma in equilibrium, the force operator is self-adjoint. By this property, Kadomtsev showed that the energy of a virtual displacement of mode  $m = 0$  was bounded in time provided that the plasma pressure gradient satisfied

$$-\frac{d \ln p}{d \ln r} \leq \frac{4\gamma}{2 + \gamma\beta} \quad (13)$$

and that this is both necessary and sufficient for stability of the static Z-pinch [1]. The condition may be recast in thermodynamic variables by considering the radial force balance

$$\frac{dp}{dr} = -\frac{B_\theta}{\mu_0 r} \frac{d}{dr}(r B_\theta) = -\frac{1}{r^2} \frac{d}{dr}(r^2 p_B). \quad (14)$$

Next consider the thermodynamic identity  $dh = v dp + T ds$ , with  $v$  specific volume, written in the form

$$\frac{dh}{dr} = \frac{1}{\rho} \frac{dp}{dr} + T \frac{ds}{dr} \quad (15)$$

where  $h = (\gamma/(\gamma-1))(p/\rho)$  is the specific enthalpy and  $s = c_v \ln(p/\rho^\gamma)$  is the specific entropy with specific heat at constant volume  $c_v = R/(\gamma-1)$ . Note that in all what follows,  $T$  signifies the total plasma temperature  $T = T_i + T_e$ , since the distinction between the two temperatures is beyond the scope of ideal MHD. Equation (15) is an algebraic identity between the change in enthalpy, the  $p dv$  work, and the reversible non- $p dv$  work expressed as  $T \nabla s$ . In this sense as an identity, an analogous identity holds for the magnetic body force

$$\frac{d}{dr} \left( 2 \frac{p_B}{\rho} \right) = \frac{1}{\rho r^2} \frac{d}{dr}(r^2 p_B) + \frac{p_B}{\rho} \frac{d}{dr} \ln \left( \frac{p_B}{\rho^2 r^2} \right). \quad (16)$$

We recognize the specific magnetic enthalpy as  $2p_B/\rho$  and the specific Laplace body force as  $(\vec{j} \times \vec{B})_r/\rho$ , leading us to label

the quantities  $T_m \equiv p_B/(\rho R)$  as a ‘‘magnetic temperature’’ and  $s_m \equiv R \ln(p_B/(\rho^2 r^2))$  as a ‘‘magnetic entropy.’’ In fact,  $s_m$  is the specific magnetic flux (or frozen-in flux) of Alfvén’s law applied to the axisymmetric pinch

$$\frac{d}{dt} \left( \frac{B_\theta}{\rho r} \right) = 0. \quad (17)$$

It is distinct from both Pastukhov’s notion of magnetic entropy in [2] and Minardi’s magnetic entropy concept based in information theory [38]. However, it is clear that in magnetohydrodynamics, the specific flux  $B/\rho$  plays the role of an entropic coordinate in the sense of (16). Furthermore, summation of (15) and (16) gives the thermodynamic law for the magnetofluid

$$\begin{aligned} \frac{d}{dr} (c_p T + 2RT_m) - \frac{1}{\rho} \frac{d}{dr} (p + p_B) \\ = T \frac{ds}{dr} + T_m \frac{d}{dr} \left( R \ln \left( \frac{p_B}{\rho^2} \right) \right) \end{aligned} \quad (18)$$

as if ‘‘ $\gamma_m = 2$ ,’’ a useful mnemonic. It seems appropriate to point out here that the specific flux  $B/\rho$  plays a double role as an entropic thermodynamic coordinate in the sense of (18) and as a vorticity in 3-D flow dynamics through the generalized (or canonical) vorticity [39].

Now with the variables  $s$  and  $s_m$ , we eliminate pressure from (14) in favor of the variables  $(\rho, s, s_m)$

$$p = \rho^\gamma e^{s/c_v} \quad (19)$$

$$p_B = \rho^2 r^2 e^{s_m/R} \quad (20)$$

and after some algebra obtain a differential equation for  $\rho$

$$c^2 \frac{d\rho}{dr} = -4 \frac{p_B}{r} - \frac{1}{c_v} p \frac{ds}{dr} - p_B \frac{ds_m}{dr} \quad (21)$$

where  $c^2 = c_s^2 + c_a^2$  is the local magnetosonic speed. Equation (21) is analogous to (4) with the difference that the body force  $\nabla \Phi_g$  is due to the geometric term  $4p_B/r$  of field line curvature and there are two thermodynamic variables  $(s, s_m)$ . Then dividing by magnetic pressure and replacing the density gradient by the gradient of pressure, we obtain a form of the static equilibrium as (with  $c_p = \gamma c_v$ )

$$K(r) \equiv \frac{d \ln p}{d \ln r} + \frac{4\gamma}{2 + \gamma\beta} = \frac{2}{2 + \gamma\beta} \frac{d}{d \ln r} \left( \frac{s}{c_p} - \frac{s_m}{2R} \right). \quad (22)$$

We recognize the left-hand side as Kadomtsev’s stability function  $K(r)$ . Thus, the right-hand side is equivalent in equilibrium. Furthermore, the logarithms combine to define  $s_z$  as

$$\frac{s_z}{R} \equiv \frac{s}{c_p} - \frac{s_m}{2R} = \ln \left( \frac{p}{p_B^{\gamma/2} r^\gamma} \right) \quad (23)$$

so we write the marginal condition  $K = 0$  as

$$\frac{ds_z}{dr} = 0. \quad (24)$$

In this way, the Z-pinch is stably stratified to axisymmetric modes through (24) in a similar way as the stably stratified atmosphere, and the development of weak instability acts to

bring the plasma toward this marginal state, as in the quasilinear theory of [40]. For example, for the axisymmetric-stable Bennett pinch of  $\gamma = 2$ , the invariant  $s_z$  defines the Z-pinch adiabat through  $\beta r^2 = e^{s_z}$ , and in the incompressible, isothermal limit  $\gamma = 1$ , one has  $\beta I_{\text{enc}} = e^{s_z}$  where  $I_{\text{enc}}$  is the axial current enclosed at radius  $r$ . In general, though, this relationship can be expressed as the statement that  $p \sim (B_\theta/r)^\gamma \sim (I_{\text{enc}}/\pi r^2)^\gamma$ , with equality given some constants which absorb into the adiabatic constant  $e^{s_z}$ . The quantity  $I_{\text{enc}}/\pi r^2 \equiv \langle j_z \rangle$  is in fact the area-averaged current density, so the marginal state reflecting a vanishing gradient of  $s_z$  has

$$\frac{p}{p_0} = \left( \frac{\langle j_z \rangle}{j_0} \right)^\gamma \quad (25)$$

which is evocative of the usual adiabatic condition  $p = \rho^\gamma$ , although the condition is nonlocal as the local value of pressure depends on the distribution of current within the plasma column up to that radius for  $p = p(r)$ . In analogy to meteorology, we refer to a pinch whose pressure exceeds the polytropic relationship of (25) as super-magnetoadiabatic, in which case it is convectively unstable to magnetic flux interchange, and the opposite case as sub-magnetoadiabatic. In this terminology, Kadomtsev’s pinch is the magnetoadiabatic pinch. Just as in meteorology, this terminology is simple but potentially misleading as the adiabatic atmosphere is not ‘‘adiabatic,’’ but rather isentropic, and the magnetoadiabatic pinch is also not ‘‘adiabatic,’’ but rather the specific magnetic flux  $s_m$  is tied to the specific entropy  $s$  such that  $\nabla s_z = 0$ .

In other words, (23) is a single equation in the two variables  $s$  and  $s_m$ , and thus any entropy profile  $s = s(r)$  such that  $\nabla s_m = 2(\gamma - 1)/\gamma \nabla s$  represents a Kadomtsev pinch. This freedom is what is usually understood to be the freedom in the choice of temperature and density profiles for the equilibrium pinch. Two situations of note are the adiabatic pinch, where  $\nabla s = 0$ , and the isothermal pinch where  $p/\rho = \text{const}$ . In the former situation, the adiabatic pinch has both  $\nabla s = 0$ , and thus,  $\nabla s_m = 0$ , so that density varies radially as  $\rho \sim \langle j_z \rangle$ . In the isothermal pinch, on the other hand, one finds that  $\rho \sim \langle j_z \rangle^\gamma$ . Both the cases are magnetoadiabatic and therefore marginally stable to interchange.

In the context of the fusion deuterium-tritium (DT) neutron yield, for temperatures up to approximately 5 keV,  $\dot{Y} \sim n^2 T^4 \sim p^2 T^2$ , such that for relatively low  $T$  more concentrated temperatures are advantageous for the same pressure or supplied current, in which case the adiabatic pinch temperature profile appears to be advantageous. On the other hand, around 50 keV the reaction rate becomes independent of temperature, making more concentrated density necessary for high-yield fusion conditions. Indeed, high densities are required for  $Q > 1$  conditions in a sheared-flow-stabilized Z-pinch [41], in which case the isothermal profile is more desirable. However, nonideal instabilities known as entropy modes grow to relax  $\nabla s \rightarrow 0$ , making such instabilities important to model  $Q > 1$  Z-pinches.

### C. Entropy Modes

Section III-B found that the axisymmetric interchange instability drives the pinch toward a marginal state where the

specific flux gradient  $\nabla s_m$  is tied to the specific entropy gradient  $\nabla s$  through the interchange marginal stability condition. That the specific entropy is unconstrained, and consequently also the concentration of temperature suggests that some critical physics may be missing from this picture of convective relaxation. Kadomtsev in fact also determined this condition, building on work by a collaborator Y.A. Tserkovnikov, to be a type of drift instability caused by the oppositely directed diamagnetic heat fluxes of electrons and ions in the inhomogeneous plasma pinch [42].

Furthermore, Kadomtsev's article anticipated that such drift modes, which leave thermal pressure unperturbed to first order with varying temperature and density, were in fact entropy fluctuations working to relax the temperature gradient, and hence are termed entropy modes. Unfortunately, the dependence on  $\gamma$  was suppressed in Kadomtsev's article in favor of the numerical factor of 5/3. The important article on numerical entropy mode modeling by Angus et al. [43] restored the  $\gamma$ -dependence and gives the entropy mode stability condition

$$\frac{d \ln T}{d \ln r} \leq 1 + \left( \frac{2\gamma - 1}{2\gamma} + \frac{\beta}{4} \right) \frac{d \ln p}{d \ln r}. \quad (26)$$

Noting that the specific entropy increment is given in terms of logarithmic temperature and pressure increments by

$$\frac{1}{R} ds = \frac{\gamma}{\gamma - 1} d \ln T - d \ln p \quad (27)$$

after some simple algebra one can show that (26) combined with (13) simplifies to the condition

$$\frac{1}{c_p} \frac{ds}{d \ln r} \leq \frac{2 + \gamma\beta}{4\gamma} K(r) \implies \frac{1}{c_v} \nabla s \leq \frac{1}{2} \nabla s_z. \quad (28)$$

When relaxed to the marginally stable state, the right-hand side is identically zero from (22), giving  $\nabla s = 0$  for the specific entropy gradient. In summary, the Z-pinch entropy modes can be understood to relax the gradient of specific entropy, which combined with the interchange mode activity yields simple estimates for the fully relaxed profile as  $\nabla s = \nabla s_m = 0$ .

#### D. Z-Pinch Buoyancy Frequency

As pointed out in [42] and [43], the radial displacement frequency of a Z-pinch plasma parcel is found by linear analysis assuming purely radial displacement to be

$$\omega^2 = \frac{2c_s^2}{r^2} K(r) \quad (29)$$

with  $K(r)$  the Kadomtsev stability function. It would then appear that the thermal sound speed is principally responsible for the growth of perturbations. However, using (22) for  $\nabla s_z$ , the sonic speed  $c_s^2 = \gamma p / \rho$ , Alfvén speed  $c_a^2 = 2p_B / \rho$ , and magnetosonic speed  $c^2 = c_s^2 + c_a^2$  gives a form more akin to the usual Brunt-Väisälä frequency in cylindrical coordinates

$$\omega^2 = \frac{2}{r} \frac{c_a^2 c_s^2}{c^2} \frac{d}{dr} \left( \frac{s_z}{R} \right). \quad (30)$$

Thus, the speed appears as a combination of both the characteristic velocities. To bring out the character of this quantity a bit, we can consider the case of an isothermal Bennett

pinch with  $\gamma = 2$ , where  $c_a^2 = 2(r/r_p)^2(kT/m)$  and  $c^2 = \gamma(kT/m)(1 + (2/\gamma\beta))$

$$\omega_B^2 \approx \frac{r/r_p}{1 + (r/r_p)^2} c_s^2 \frac{d}{dr} \left( \frac{s_z}{R} \right) \quad (31)$$

such that the frequency of interchange is potentially greatest at the pinch radius  $r = r_p$ . Note that (30) is not singular as  $c_a^2$  always grows at least quadratically for center-peaked currents.

#### E. Marginal Pressure Distribution

Substituting the invariant  $s_z$  into the pressure balance and integrating produces an energy relationship

$$\frac{\gamma}{\gamma - 1} \left( \frac{p}{p_0} \right)^{(\gamma-1)/\gamma} + 2 \left( \frac{E_I}{P_0} \right)^{1/2} = \frac{\gamma}{\gamma - 1} \quad (32)$$

with  $E_I = (\mu_0/8\pi)I_{\text{enc}}^2$  and  $P_0 = p_0\pi r_p^2$ . Each term represents fluid or magnetic-specific enthalpy under adiabatic or flux-conserving conditions. That is, let the enthalpy of a plasma parcel isentropically brought to its reference pressure  $p_0$  be the potential enthalpy  $h_f^0 \equiv c_p T(p_0/p)^{(\gamma-1)/\gamma}$ , and let the potential magnetic enthalpy  $h_m^0 \equiv 2RT_m(P_0/E_I)^{1/2}$  be the magnetic enthalpy when brought to the reference magnetic pressure under constant specific flux. Then (32) is

$$\frac{h_f}{h_f^0} + \frac{h_m}{h_m^0} = \frac{c_p}{R}. \quad (33)$$

Equation (33) has the significance that specific enthalpy is radially uniform for the adiabatic ( $\nabla s = 0$ ) pinch, and otherwise varies according to the particular entropy gradient of the profile.

Now using the marginal stability condition as a differential equation between  $\beta$  and  $r$ , we observe that  $\beta E_I^{1/2} = (1 + 2(\gamma - 1)/\gamma\beta)^{-1}$ , so that (32) rearranges to

$$\frac{p_0}{p} = \left( 1 + \frac{\gamma - 1}{\gamma} \frac{2}{\beta} \right)^{\gamma/(\gamma-1)}. \quad (34)$$

To reveal the physical character of the distribution, consider the limit  $\gamma \rightarrow 1$  where the pressure varies as  $p = p_0 e^{-2/\beta}$ , or

$$p = p_0 \exp\left(-\frac{m v_a^2}{k_B T}\right) \quad (35)$$

where  $v_a$  is the Alfvén speed and  $T$  is the total temperature. In other words, the pressure is a frozen Boltzmann distribution in terms of relative magnetic-to-thermal energy analogous to the isentropic relationship for compressible flow. Then substituting (34) into the invariant  $s_z$ , one obtains

$$\left( \frac{r}{r_p} \right)^2 = \beta^{-1} \left( 1 + \frac{\gamma - 1}{\gamma} \frac{2}{\beta} \right)^{(2-\gamma)/(\gamma-1)}. \quad (36)$$

This relationship is an implicit relationship for  $\beta = \beta(r)$  to calculate the pressure, magnetic field, and current density profiles of the Kadomtsev-stable pinch.

### F. Special Cases of the Marginal Distribution

A few interesting cases present themselves for (34) and (36). The radius- $\beta$  relationship is explicitly invertible for three cases of interest. Of course, the case  $\gamma = 2$  obtains the Bennett profile where  $(r_p/r)^2 = \beta$ . There is also the limit  $\gamma \rightarrow 1$  where  $(r_p/r)^2 = \beta e^{-2/\beta}$  which inverts using the Lambert W function. In the particular case of  $\gamma = 5/3$ , (36) is of the depressed cubic form

$$\beta^3 - \left(\frac{r_p}{r}\right)^4 \beta - \frac{4}{5} \left(\frac{r_p}{r}\right)^4 = 0 \quad (37)$$

and thus possesses reasonably explicable roots for  $\beta = \beta(r)$ . A compact, explicit form particularly suitable for use in codes comes from Viète's form of the cubic root

$$\beta = \left(\frac{r_p}{r}\right)^2 \frac{2}{\sqrt{3}} \cos\left(\frac{1}{3} \arccos\left(\frac{6\sqrt{3}}{5} \left(\frac{r}{r_p}\right)^2\right)\right) \quad (38)$$

considered as a complex-valued function returning a real-valued solution, which excludes the other complex-valued and negative-valued cubic roots. From these considerations, evidently local  $\beta$  at the pinch edge radius depends on  $\gamma$ . For  $\gamma = 1$  at the pinch radius  $r_p$ , the solution to  $2/\beta = e^{-1/\beta}$  is  $\beta(r_p) = 2/W(2) \approx 2.34$  where  $W(r)$  is Lambert's W function. If the pinch radius was rescaled by  $\sqrt{2}$  (as suggested by the relationship  $r_p^2 = \lambda_D^2/4(c/v_e)$  for the Bennett pinch [44]), then the pinch-edge beta satisfies  $\beta(r_p/\sqrt{2})/2 = \Omega^{-1} \approx 1.763$  where  $\Omega$  is the Omega constant of the Lambert W function. Transcendentals aside, evidently there is a tendency toward greater  $\beta$  as  $\gamma \rightarrow 1$ , and this is accomplished at the expense of a larger specific heat for compression. Fig. 1 shows the variation in pinch-edge radius  $\beta$  with adiabatic index  $\gamma$ .

In the infinite degree-of-freedom limit  $\gamma \rightarrow 1$ , the pressure has a simple explicit form in terms of the Lambert W function. With  $2/\beta = W(2(r/r_p)^2)$ , the Kadomtsev-stable pressure is

$$p(r) = p_0 \exp(-W(2(r/r_p)^2)) \quad (39)$$

and the profile for other values of  $\gamma \in (1, 2]$  can be understood to limit between this form and the Bennett pinch whose explicit pressure profile is  $p(r) = p_0/(1 + (r/r_p)^2)^2$ . From this, we note that the pinch-edge pressure is  $p = p_0/4$  for the Bennett pinch and  $p = W(2)/2p_0 \approx 0.426p_0$  for the  $\gamma \rightarrow 1$  limit. When pinch radius is rescaled by  $\sqrt{2}$ , then at  $r = r_p/\sqrt{2}$ , the pressure  $p = \Omega p_0$  and the magnetic pressure  $p_B = \Omega^2 p_0$ . We also note another application of the invariant (32) as an implicit relationship for  $p = p(r)$  valid for all  $\gamma$  as

$$\frac{\gamma}{\gamma - 1} \left(\frac{p}{p_0}\right)^{(\gamma-1)/\gamma} + 2 \left(\frac{r}{r_p}\right)^2 \left(\frac{p}{p_0}\right)^{1/\gamma} = \frac{\gamma}{\gamma - 1}. \quad (40)$$

### G. Pinch Radius for Marginally Stable Profiles

In this section, we demonstrate how the  $\beta = \beta(r)$  relationship expressing the distribution of the magnetic energy per thermal energy is closely connected to the general result for the inductance of a current-carrying pinch, and in this way we determine how the pinch's radius is related to its total thermal and magnetic energies. To begin, recall that Z-pinches confine

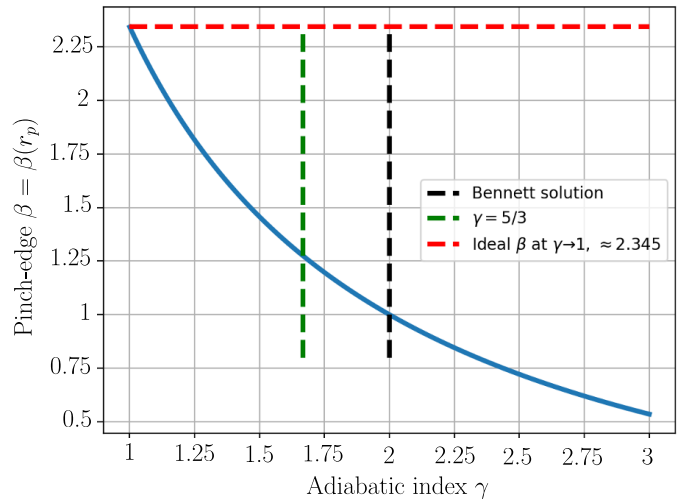


Fig. 1. Local plasma  $\beta = p/p_B$  at the characteristic radius  $r_p$  (i.e., the pinch-edge beta) of the Kadomtsev-stable pinch increases as the adiabatic index  $\gamma \rightarrow 1$ . For the Bennett pinch,  $\beta = 1$  at the pinch radius, and for  $\gamma \in (1, 2]$  the pinch-edge  $\beta$  takes an intermediate value.

plasma with unity average  $\beta$  [45]. Precisely, the Bennett relationship for self-magnetic confinement relates the self-magnetic energy of axial current  $(\mu_0/8\pi)I_{\text{enc}}^2$  to the excess thermal pressure contained within an area of cross-sectional radius  $r$

$$\frac{\mu_0}{8\pi} I_{\text{enc}}^2(r) = \pi r^2 (\langle p \rangle - p(r)) \quad (41)$$

where  $\langle p \rangle \equiv (\pi r^2)^{-1} \int_0^r p(r') 2\pi r' dr'$  is the area-averaged thermal pressure. Out to a radius where  $p \rightarrow 0$  (41) gives

$$\frac{\langle p \rangle}{p_B(r)} = 1 \quad (42)$$

with  $p_B(r)$  the magnetic pressure at that radius. In this sense, the average Z-pinch  $\beta$  is unity, and indeed nonzero axial magnetic field results in a value less than unity. Essentially, (41) has the significance that purely axial current ideally confines an excess of thermal energy in an amount equal to the current's self-magnetic energy [46].

Of course, in addition to the self-magnetic energy of the internal current, the total magnetic energy is fixed by the radius of the return current as in a coaxial cable. In the context of the internal conductor as a Z-pinch, with fixed outer radius and variable inner radius, this is related to the spatial localization of the confined thermal energy (i.e., the internal conductor radius). By combining the physics for the inductance of a coaxial conduit with the Bennett relationship, we find that the Z-pinch radius is a function of a quantity for which both the numerator and the denominator of (42) are averaged.

To demonstrate, we analyze the sharp pinch for simplicity and then state results for some special cases of the Kadomtsev-stable profiles. Consider a sharp pinch of radius  $r_p$  contained in a perfectly conducting wall of radius  $r_w$  returning all the current passed by the sharp pinch. In this situation, the profiles

of thermal and magnetic pressures are given by

$$p = \begin{cases} p_0, & r < r_p \\ 0, & r_p < r < r_w \\ 0 & r > r_w \end{cases}, \quad p_B = \begin{cases} 0, & r < r_p \\ p_0 \left(\frac{r_p}{r}\right)^2, & r_p < r < r_w \\ 0 & r > r_w. \end{cases} \quad (43)$$

Now calculating the total thermal and magnetic energies per unit length by integrating over all radii, we find a thermal energy per unit length in the amount of  $P_T = p_0 \pi r_p^2$  and likewise magnetic energy per unit length of  $P_B = p_0 \pi r_p^2 \ln((r_w/r_p)^2)$ . Now defining the ratio of linear pressure to magnetic pressure as  $\beta_m \equiv P_T/P_B$  (i.e., (42) with both the numerator and denominator area-averaged), one has

$$\left(\frac{r_p}{r_w}\right)^2 = e^{-1/\beta_m}. \quad (44)$$

Indeed, it is interesting to observe that the isothermal sharp pinch size is Boltzmann distributed as  $\exp(-((P_B/N)/k_B T))$  with  $N$  the linear density. The quantity  $\beta_m$  measures the spatial concentration of pinch energy according to the usual physics of a coaxial conductor combined with the Bennett relationship. Consequently, to localize the plasma in a small radius, the parameter  $\beta_m \ll 1$ . For a fixed linear density  $N = n_0 \pi r_p^2$  with  $n_0$  the volumetric particle density, high density requires a small pinch radius  $r_p$ , and high density pinch discharges are required for  $Q > 1$  conditions [41]. Therefore, the total magnetic energy of a  $Q > 1$  Z-pinch discharge must be large compared with the confined thermal energy. However, we emphasize that this result occurs precisely because the Z-pinch has unity average- $\beta$ . That is, the local  $\beta$  at the pinch edge is approximately unity, and most of the magnetic energy occurring in such a low- $\beta_m$  Z-pinch discharge is localized as either vacuum field or in a low- $\beta$  edge plasma. Indeed, for precisely this reason magnetically confined plasmas with axial magnetic fields such that  $\beta$  is low everywhere consequently require even greater quantities of magnetic energy for thermal confinement.

The result  $r_p^2 \sim e^{-1/\beta_m}$  holds asymptotically for the Kadomtsev-stable Z-pinch profiles. Lacking a return current, the total magnetic energy is divergent for all the values of  $\gamma$ . However, the magnetic energy is small enough for  $r \gg r_p$  that we may cut off the integral at the wall for a sufficient approximation. In the case of the Bennett pinch, an exact solution can be found, but since the cutoff is not valid for  $r_p \approx r_w$ , we approximate  $r_p \ll r_w$  for

$$\left(\frac{r_p}{r_w}\right)^2 \approx e^{-1} e^{-1/\beta_m} \quad (45)$$

such that the diffuse Bennett pinch has a smaller radius than a sharp pinch of equivalent thermal and magnetic energies, being one  $e$ -folding factor smaller for the same value of  $\beta_m$ . A similar calculation of the Kadomtsev pinch for  $\gamma = 5/3$  demonstrates the equivalent asymptotic behavior as a function of  $\beta_m$  with an additional benefit of approximately one more  $e$ -folding factor, as shown in Fig. 2. The physical reason for the asymptotic variation as  $e^{-1/\beta_m}$  is the inductive coupling of the current contained within  $r \leq r_p$  and the return

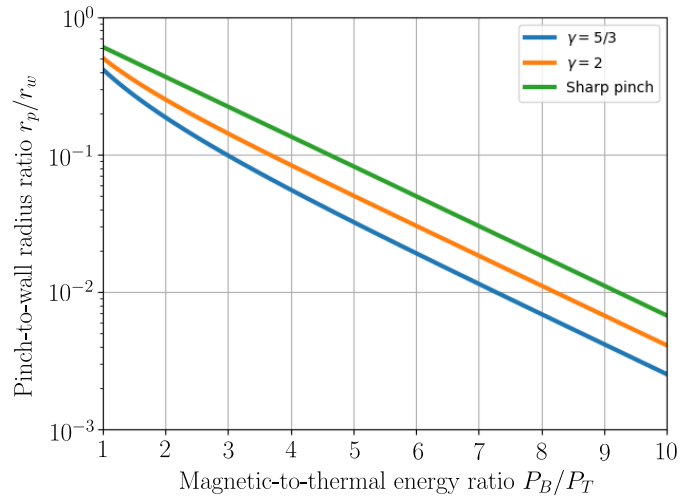


Fig. 2. Z-pinch size asymptotically varies exponentially with the ratio of total confined thermal to magnetic pressures. Here,  $P_B$  is the magnetic energy per unit length and  $P_T$  the pressure energy per unit length. For the Kadomtsev-stable profiles, smaller values of  $\gamma$  give an increasing benefit over sharp pinches in terms of confinement, but at the expense of a lower temperature for the same confined pressure as the specific heat increases.

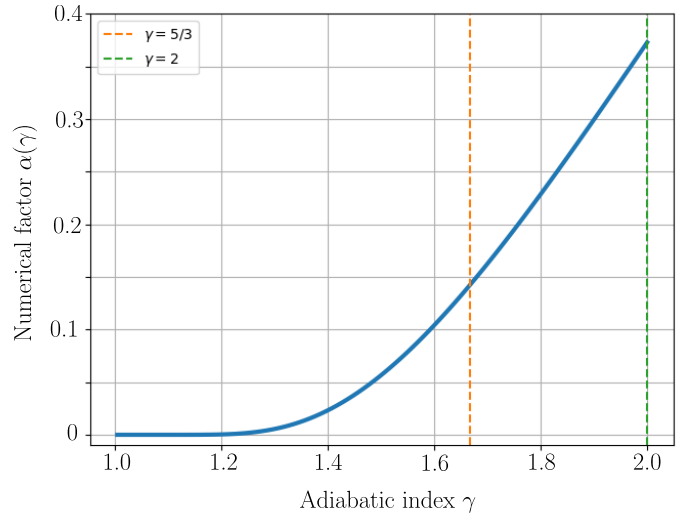


Fig. 3. Functional dependence on adiabatic index for marginally Kadomtsev-stable pinch radius according to  $(r_p/r_w)^2 = \alpha(\gamma)e^{-1/\beta_m}$  calculated numerically. The stable pinch radius collapses as  $\gamma \rightarrow 1$  for the same magnetic and pressure energies  $P_B, P_T$ , or equivalently at fixed  $\beta_m$ .

current at  $r = r_w$ , such that the inductance for  $r_p \ll r_w$  is well-approximated by the usual coax factor  $\ln(r_w/r_p)$  for  $\beta_m \lesssim 1$ .

For Kadomtsev's pinch, radius depends on adiabatic index  $\gamma$  and the parameter  $\beta_m$  through  $(r_p/r_w)^2 = \alpha(\gamma)e^{-1/\beta_m}$  (valid for  $r_p \ll r_w$ ), where  $\alpha(\gamma)$  is plotted in Fig. 3. Indeed, the total pressure energy per unit length of the Kadomtsev profile may be calculated as  $P_T = p_0 \pi r_p^2 (\gamma/(2(\gamma-1)))^2$ . From this, one can see that given fixed quantities of pressure energy  $P_T$  and magnetic energy  $P_B$ , as  $\gamma \rightarrow 1$  the pinch radius  $r_p \rightarrow 0$  and the on-axis pressure  $p_0 \rightarrow \infty$  such that  $p(r) \rightarrow P_T \delta(r)$ .

#### H. Inductance, Enthalpy, and Flux of Kadomtsev's Pinches

Section III-G introduced the energy measure  $\beta_m \equiv P_T/P_B$  and its relationship to the characteristic pinch radius. Here, the

relationship to inductance is made precise, and along the way a remarkable fact is noted that total pinch enthalpy  $H = I_\infty \psi'$  where  $\psi'$  is the flux per unit length and  $I_\infty$  the total enclosed current.

We begin by calculating the magnetic flux per unit length  $\psi' = \int_0^r B_\theta dr$  (or magnetic vector potential  $A_z = -\psi'$ ) as

$$\frac{\psi'}{\psi'_0} = \frac{\gamma}{2\sqrt{2}(\gamma-1)^2} \left( \log(1+\Lambda) + \frac{\gamma-2}{1+\Lambda^{-1}} \right) \quad (46)$$

from  $\psi' = \int_\beta^\infty B_\theta(\beta')(dr/d\beta')d\beta'$ , where  $\Lambda \equiv 2(\gamma-1)/(\gamma\beta)$  and  $\psi'_0 = B_0 r_p$  with  $B_0 = \sqrt{\mu_0 p_0}$  the characteristic flux density. Then computing the total enthalpy obtains (with  $P_0 = p_0 \pi r_p^2$ )

$$H = \frac{\gamma}{\gamma-1} P_T + 2P_B = \sqrt{2} \frac{\gamma}{\gamma-1} P_0 \frac{\psi'}{\psi'_0}. \quad (47)$$

From  $\mu_0 I_\infty^2 / 8\pi = (\gamma/(2(\gamma-1)))^2 P_0$ , we have  $H = I_\infty \psi'$ . Here,  $H$  and  $\psi'$  are integrated up to a cutoff while  $I_\infty$  is the total current. Notably, under uniform axial flow  $v_z$ , enthalpy flux is precisely circuit power  $H v_z = \phi I_\infty$  where voltage  $\phi = v_z \psi'$ . Finally, as  $\psi' = L'I$  with  $L'$  inductance per unit length, dividing (47) and applying the Bennett relationship gives

$$L' = \frac{\mu_0}{4\pi} \frac{\gamma}{2(\gamma-1)} \left( 1 + \frac{2(\gamma-1)}{\gamma} \beta_m^{-1} \right) \frac{\Lambda}{1+\Lambda}. \quad (48)$$

Equation (48) is well-approximated by  $\Lambda/(1+\Lambda) \approx 1$  for cutoff  $\beta \ll 1$ . For example, the  $\gamma = 2$  Bennett pinch inductance is  $L' \approx (\mu_0/4\pi)(1 + \beta_m^{-1})$ . Notable is the  $\lim_{\gamma \rightarrow 1} L' = (\mu_0/4\pi)\beta^{-1}$ .

#### IV. MARGINAL PROFILE AND SHEARED AXIAL FLOWS

This section begins with some theoretical remarks on sheared axial flows under completely ideal conditions in Z-pinch experiments, and then presents some observations from resistive MHD modeling of the FuZE experiment [45], [47]. In the sheared-flow-stabilized Z-pinch concept, a flow Z-pinch is formed downstream from a coaxial plasma gun. As a starting point, we draw some basic conclusions about steady flow using MHD theory in a similar manner to the analysis done by Morozov and Solov'ev [48] and Morozov [49]. To begin, recall that in compressible ideal MHD there are three axisymmetric streamline invariants, namely, specific enthalpy, specific entropy, and specific magnetic flux

$$h_t \equiv \frac{\gamma}{\gamma-1} \frac{p}{\rho} + 2 \frac{p_B}{\rho} + \frac{v^2}{2} \quad (49)$$

$$s \equiv \frac{R}{\gamma-1} \ln \left( \frac{p}{\rho^\gamma} \right) \quad (50)$$

$$s_m \equiv R \ln \left( \frac{p_B}{\rho^2 r^2} \right). \quad (51)$$

By expressing the convective derivative  $(\vec{v} \cdot \nabla)\vec{v}$  in terms of its potential and vortex parts, we may also formulate Crocco's theorem in steady state [50]

$$\vec{v} \times \vec{\omega} = \nabla h_t - T \nabla s - T_m \nabla s_m \quad (52)$$

where  $\vec{\omega} = \nabla \times \vec{v}$ . If we consider purely axial downstream flow such that flow forces are zero, the force balance is

$$\frac{dh}{dr} = T \frac{ds}{dr} + T_m \frac{ds_m}{dr} \quad (53)$$

where the specific MHD enthalpy  $h = h_t - v^2/2$  does not include the kinetic energy. If the flow is self-organized into a near-Kadomtsev-stable profile with  $\nabla s_z = 0$ , then

$$\frac{dh}{dr} = \Theta \frac{ds}{dr} \quad (54)$$

where  $\Theta = T(p_0/p)^{(\gamma-1)/\gamma}$  is the potential temperature (which is only constant radially for the adiabatic pinch with  $\nabla s = 0$ ). Equation (54) follows by factoring  $T \nabla s$  from (53) and eliminating  $\beta = T/T_m$  using (34). Expressing the specific entropy as  $s = c_p \ln \Theta$  shows that (54) integrates into the Kadomtsev pinch's specific enthalpy profile [which should be distinguished from the invariant of (33)]

$$h = c_p \Theta = c_p T \left( 1 + \frac{\gamma-1}{\gamma} \frac{2}{\beta} \right). \quad (55)$$

As the specific enthalpy is conserved along streamlines, a first approximation is an isoenergetic condition transverse to the flow, i.e.,  $(dh_t/dr) = 0$ , supposing that flow originates from one "reservoir." In the isoenergetic case, the energy relationship is

$$c_p \Theta + \frac{v^2}{2} = \text{const.} \quad (56)$$

Equation (56) means that the axial flow profile  $v_z = v_z(r)$  of a pinch self-organized into a Kadomtsev-stable state downstream from its source is ideally a function only of the specific entropy profile  $s = s(r)$  through  $\Theta(r) = \Theta_0 e^{s(r)/c_p}$ . Because the pressure profile  $p = p(r)$  is fixed by marginal stability, this is equivalent to thinking of the flow shear  $\partial_r v_z$  as a function of the temperature profile  $T = T(r)$  in the first approximation. Furthermore, (56) predicts that when  $\nabla s = 0$  the axial flow is radially uniform. Because entropy modes drive  $\nabla s \rightarrow 0$ , this suggests that flow shear of a pinch marginally stable to both interchange and entropy modes must be controlled through engineering the gradient of specific enthalpy exiting the coaxial plasma accelerator, for example, through an oblique shock wave attached to the tip of the central electrode.

Considering isothermal pinch profiles, for a given uniform specific enthalpy  $h_t$ , say of the reservoir, (56) has no solution past some radius. For example, the Bennett pinch has  $\beta = (r_p/r)^2$ , such that the energy relationship reads like  $c_1 v^2 + c_2 r^2 = c_3$  for positive constants  $c_1$ ,  $c_2$ , and  $c_3$ , defining an elliptical velocity profile, decreasing from a maximum on axis. This result occurs because isothermal pinches have specific enthalpy  $h \rightarrow \infty$  as  $r \rightarrow \infty$ , which is problematic in experiment. Under the isoenergetic assumption, we can determine the conditions for axial velocity to be an increasing function of radius by writing (56) as

$$v^2 = 2h_{t0} \left( 1 - \frac{c_p \Theta_0}{h_{t0}} e^{s/c_p} \right) \quad (57)$$

with  $h_{t0}$  the given constant specific enthalpy. Equation (57) suggests that for velocity to increase with radius, entropy



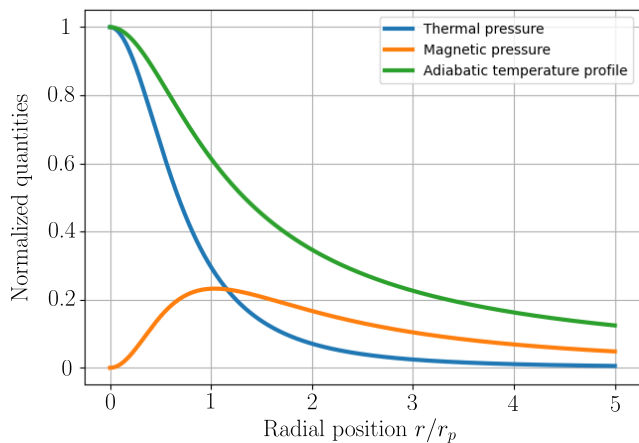


Fig. 4. Thermal and magnetic pressures of the Kadomtsev pinch profile are depicted alongside the adiabatic temperature profile (where specific entropy  $s$  is relaxed to  $\nabla s = 0$ ), with pressures normalized to  $p_0$  and the temperature normalized to  $T_0$ . The adiabatic temperature profile is marginally stable to the nonideal entropy modes for the pressure profile of Kadomtsev’s pinch. Temperature profiles which decrease more slowly than the depicted, or even increase with radius, are superadiabatic ( $\nabla s > 0$ ) and unstable to the nonideal entropy modes similar to the Schwarzschild criterion for convective stability.

should decrease with  $\nabla s \leq 0$ , which is also the entropy mode stability condition given marginal interchange stability. The temperature profiles with  $\nabla s < 0$  are called subadiabatic. Such profiles are axially peaked and decrease at least as fast as the adiabatic profile  $T(\beta) = T_0(1 + 2(\gamma - 1)/\gamma\beta)^{-1}$  depicted in Fig. 4. These considerations suggest that adiabatic temperature profiles are to be expected rather than isothermal ones. On the other hand, velocity may freely increase with radius for arbitrary profiles of specific enthalpy.

It must be cautioned here that our discussion is not self-consistent because the flow shear predicted by (56) for the marginally stable state also modifies the stability condition in a nontrivial manner, allowing for super-magnetoadiabatic pressure profiles [27] (and most importantly, kink stability and 3-D self-organized states). There is also the issue of weak collisionality; the axisymmetric minimum energy state is Kadomtsev distributed only under the ideal fluid closure. The kinetic equilibrium of a weakly collisional Z-pinch induces non-Maxwellian distribution functions in the presence of, e.g., axially sheared flow or temperature gradients, as the ion distribution function is distributed over the canonical momentum  $P_z = mv_z + qA_z$  where  $A_z$  is the magnetic vector potential [51]. This introduces intriguing phenomena such as a tendency for velocity to trend with the magnetic flux function [52], [53] which are beyond the scope of the fluid model. That is to say, simulation of the appropriate model is necessary to self-consistently model flow pinch behavior in the weakly collisional regime of the FuZE experiment.

#### A. Results From MHD Modeling of the FuZE Experiment

Here observations are presented of whole device 2-D axisymmetric resistive magnetohydrodynamic modeling of the FuZE experiment, a diagram of which is shown in Fig. 5, using the WARPXM discontinuous Galerkin finite element code [54]. The modeling approach uses an unstructured mesh of the whole device (including the accelerator) to solve the

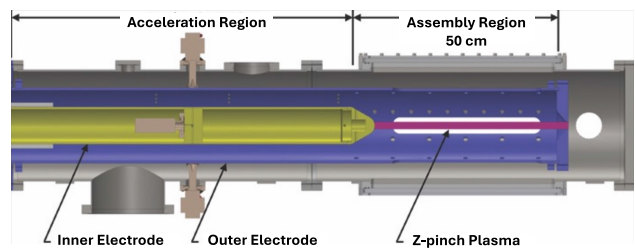


Fig. 5. Diagram of the FuZE sheared-flow-stabilized Z-pinch device, showing the coaxial plasma accelerator of 100-cm length coupled to the 50-cm-long pinch assembly region [45]. The phrase “terminus of the inner electrode” refers to the tip of the central conductor marking the end of the coaxial accelerator and the beginning of the pinch assembly region.

resistive MHD model with von Neumann–Richtmyer artificial viscosity coupled to a circuit model for the capacitor bank discharge. The simulations are initialized with plasma in the accelerator and the capacitor bank connected across the inner and outer electrodes. Simulations show plasma leaving the “acceleration region” (between the inner and outer electrodes) and forming a Z-pinch plasma in the “assembly region” (downstream from the terminus of the inner electrode). Related publications [47], [55] include the whole device simulations used for the comparisons here. Refer to [55] for all details of simulation methodology.

We investigate the role of marginal profiles in the data obtained by modeling. The results of this section are not meant to robustly model all the aspects of the sheared-flow-stabilized Z-pinch and its stability, in which 3-D, two-fluid [56], and kinetic physics play key roles, but rather to demonstrate the robustness of attraction to organized states under rapid changes in the discharge, and to suggest the structure of the flow pinch to consist of a central, close-to-marginal flowing core surrounded by a low- $\beta$  sheared flow. We consider the profiles in two representative stages pre- and postcompression to examine the profiles at peak performance.

Regarding the used model, resistive MHD simulation certainly captures the drive toward marginal interchange stability. However, only two-fluid modeling further captures the drive toward the adiabatic pinch temperature profile through activity of the entropy modes. Furthermore, as the pinches of the FuZE experiment are approaching the large Larmor radius regime, future research will consider the role of kinetic equilibrium and associated phenomena in the fluid description of weakly collisional flow Z-pinch.

1) *Diffuse Flow Pinch Prior to Compression, or “Zippering”*: Fig. 6 shows the current and voltage traces produced by experiment and simulation of a typical high-performance FuZE discharge. Considering the flow pinch properties sufficiently far downstream from the terminus of the inner electrode of the coaxial accelerator such that the flow is primarily axial (specifically, 20-cm downstream from the terminus), Fig. 7 shows the radial plasma profiles prior to pinch “zippering” (also referred to as pinch assembly, as depicted in [57, Fig. 3]). In this terminology, flow pinch zippering is similar to the classical zippering of gas puff pinch experiments [58]. The Kadomtsev pressure profile of Fig. 7 is calculated by (34) and the local  $\beta = p/p_B$  with  $\gamma = 5/3$ .

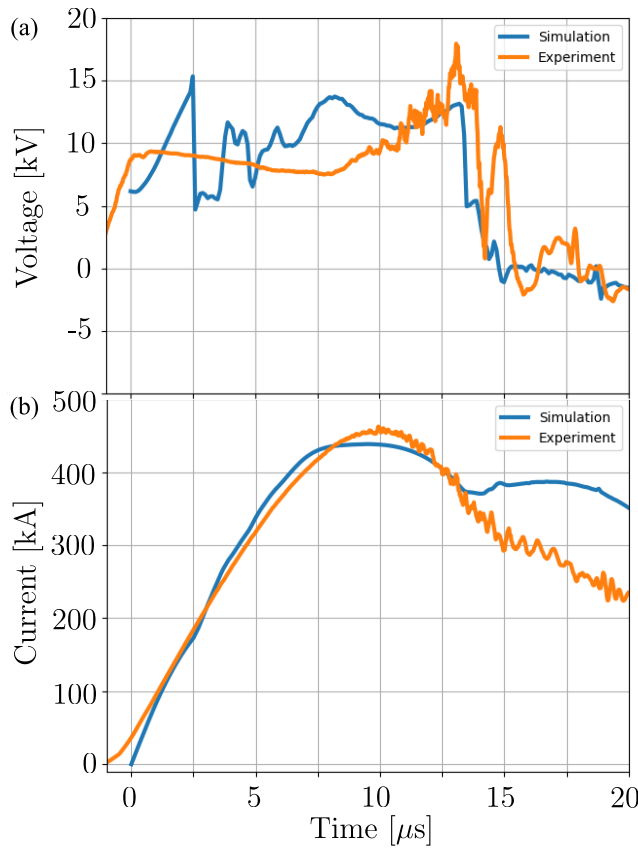


Fig. 6. Voltage (a) and current (b) traces of a 25-kV FuZE discharge compared with the same traces from 2-D resistive MHD modeling. Specifically, this voltage is measured at the gap between the inner and outer electrodes. The simulation voltage is thought to overpredict conditions around  $t = 8 \mu\text{s}$  due to a blow-by instability observed in simulation within the coaxial accelerator, which may not occur in this way under these experimental conditions, instead being modified by 3-D filamentation and the Hall effect. On the other hand, the peak voltage around  $t = 14 \mu\text{s}$  is captured well.

It is thought that the flow pinch observed in Fig. 7 is only partially representative of experimental conditions due to the limitations of our model, but it is nevertheless interesting to observe the dynamics in the context of stability theory. To interpret the results, recall that linear flow shear (in ideal MHD, of either shear orientation) is stabilizing to the convective modes. The pinch is well described by a Kadomtsev-stable profile from the axis up to approximately the pinch radius, and axial velocity decreases with radius. In this situation, the pinch core is primarily in a marginally stable equilibrium in which flow shear has not significantly altered the pressure distribution  $p = p(\beta)$  away from Kadomtsev’s pinch profile, and for which the flow shear is thought to influence stability of the kink modes (or to encourage the existence of a 3-D relaxed state). The specific enthalpy is not radially uniform so that (56) does not directly apply, and the temperature and kinetic energy are observed to trend with one another. That  $\nabla h_t \neq 0$  is thought to occur due an oblique shock upstream at the terminus of the central electrode which breaks the ideal streamline invariants including total specific enthalpy  $h_t$ . The thermal pressure profile tracks closely with the magnetoadiabatic expectation  $p \sim \langle j_z \rangle^\gamma$  once corrected by an estimate for the excess background pressure which plays no role in the pinch equilibrium.

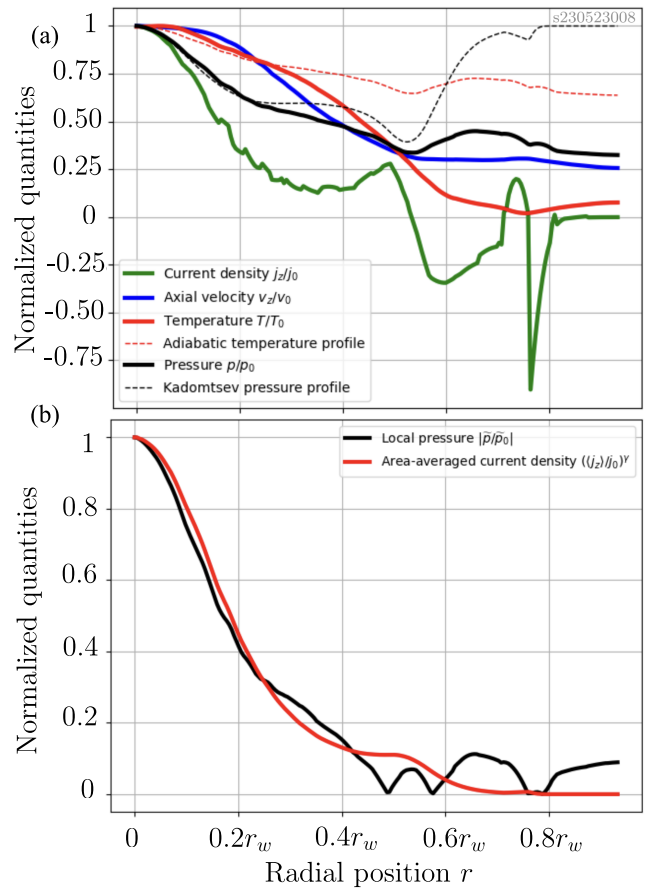


Fig. 7. Simulated pinch radial profiles prior to “zippering,” at  $t = 8 \mu\text{s}$  in the discharge, measured 20 cm downstream from the terminus of the coaxial discharge channel showing (a) radial profiles of current density, axial velocity, temperature, and pressure and (b) the corrected pressure  $|\tilde{p}|$  and area-averaged current density  $(\langle j_z \rangle)$  raised to the power of  $\gamma$  as  $p \sim \langle j_z \rangle^\gamma$ , i.e., the magnetoadiabatic profile. Corrected pressure is defined as  $\tilde{p} \equiv p - p_{\text{excess}}$  where  $p_{\text{excess}}$  is the excess pressure taken to be  $p_{\text{excess}}/p_0 = 0.38$  and must be used for magnetoadiabatic comparison because a uniform excess pressure is unknown to the MHD force balance. Here,  $r_w$  is the conducting wall of 10-cm radius. The Mach number at  $r = 0$  is  $M \approx 1.8$ . The dashed black line depicts the pressure profile if it were Kadomtsev-distributed with the local  $\beta$ , taking  $\gamma = 5/3$ . The diffuse flow pinch approximately follows Kadomtsev’s profile until a radius with significant shear where the distribution is super-magnetoadiabatic. Profile agreement ceases where the low-density edge plasma carries return current (where  $j_z < 0$ ). The flow pinch is observed to be stable for many Alfvén transit times, hence the super-magnetoadiabatic profile around  $r = 0.3r_w$  could be attributed to linear flow shear.

2) Profiles of the Post-“Zippering” Compressed Flow Pinch: Following “zippering” of the pinch discharge, the diffuse flow pinch of Fig. 7 is near-adiabatically compressed when magnetic energy increases significantly, and the pinch core shrinks to a small radius in line with Section III-G. Fig. 8 shows such a simulated post-zippering flow pinch and a comparison of its pressure to area-averaged current density. The profile within the compressed pinch radius is close to marginal in the sense that expected  $m = 0$  growth times consist of many magnetosonic transit times, as described below, and the profile is surrounded by a low- $\beta$  edge plasma with sheared axial flow deviating from the marginal state. A layer of edge plasma around  $r = 0.5r_w$ , out of view of the figure, carries some return current. A super-magnetoadiabatic profile is observed in the low- $\beta$  plasma layer in-between the pinch edge and the return current-carrying edge. Namely, an “isothermal-like”

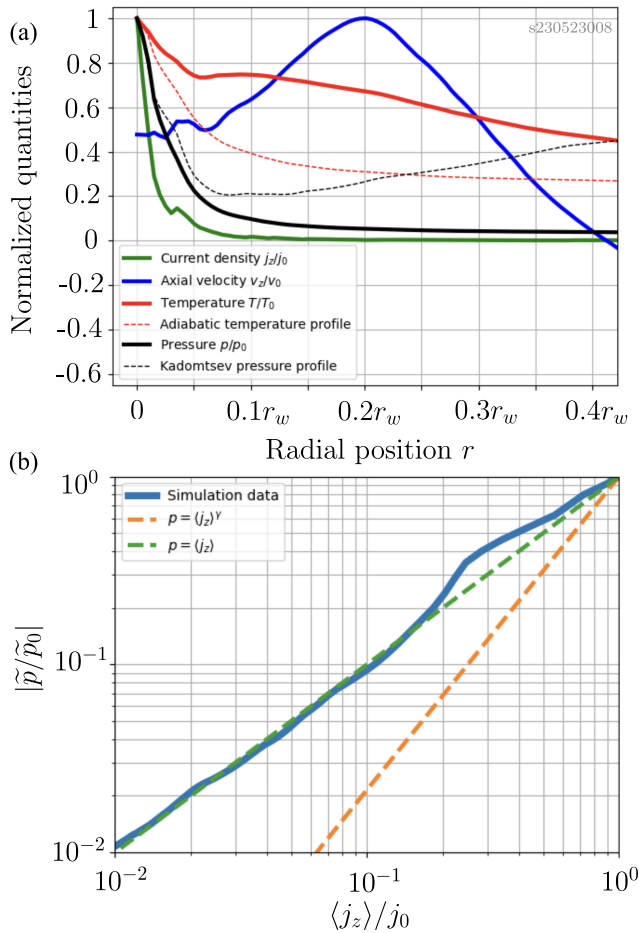


Fig. 8. Conditions at time  $t = 13 \mu\text{s}$  into the discharge, prior to the loss of flow power and just after “zippering” into small radius and maximum performance (as measured by a posteriori neutron yield rate calculation), with (a) plot of Fig. 7(a) zoomed radially up to  $r = 0.4r_w$  and (b) comparison of corrected pressure  $|\tilde{p}|$  to averaged current density  $\langle j_z \rangle$  for the data of (a), where  $p_{\text{excess}} = 0.03 p_0$  (see Fig. 7 for definition of corrected pressure). In (b), the solid curve plots  $\tilde{p} = \tilde{p}(\langle j_z \rangle)$  while the dashed lines are for reference. The profile is moderately sheared outside of the core region as the peak magnetosonic Mach number  $M_{\text{max}} = 0.2$ . A super-magnetoadiabatic profile, i.e., with better-than-Kadomtsev confinement, is observed at all radii. Outside the pinch radius  $r_p \approx 0.05 r_w$ , the profile is observed to follow the super-magnetoadiabatic scaling  $|\tilde{p}| \sim \langle j_z \rangle$ . This state is observed to last until the end of plasma voltage around  $t = 14 \mu\text{s}$  in Fig. 6, at which point the flow-sustaining electromagnetic power is extinguished. In the 2-D axisymmetric simulation, the pinch then fully relaxes into a static Kadomtsev profile, yet in 3-D would quickly undergo static kink instability.

profile  $p \sim \langle j_z \rangle$  (better than Kadomtsev) is observed in this layer rather than the magnetoadiabatic profile  $p \sim \langle j_z \rangle^\gamma$ . A clear interpretation of this observation is difficult, but we do note that every polytropic relation  $p \sim \langle j_z \rangle^a$  corresponds with some polytropic MHD equilibrium.

To clarify the phrase “close to marginal” used above, consider that if the profile were strongly unstable we could estimate from the displacement frequency of (29) a growth-time  $\tau \approx r_p/c$  where  $c$  measures the magnetosonic speed (with  $K \approx 1$  at the pinch edge for a strongly unstable profile). Using  $r_p \approx 0.5 \text{ cm}$  and  $c \approx 6.5 \times 10^5 \text{ m/s}$ , we find a growth-time of  $\tau \approx 10 \text{ ns}$ , which is not observed. However, using the observed scaling  $p \sim \langle j_z \rangle$ , we estimate  $\tau^{-2} \approx 2(\gamma - 1)(c^2/r_p)(-\nabla \ln p)$  from (30) and obtain  $\tau \approx 3.5 \mu\text{s}$ , which is comparable to the steady flow-through and dynamical timescales.

In addition to the single simulation examined here, we have considered simulations conducted under a variety of alternative realizations of discharge voltage, current, and capacitor bank configuration. Conditions similar to the results discussed previously are observed in each realization. Namely, pinch pressure and current are, prior to “zippering,” Kadomtsev-distributed (i.e., following the polytropic relationship  $p \sim \langle j_z \rangle^\gamma$ ) with  $\beta_m \lesssim 1$ . The pressure distribution breaks from marginal stability following compression, yet is often observed to remain polytropic with an exponent less than  $\gamma$ . We leave a thorough investigation of this phenomenon to future work.

To summarize, we observe that axisymmetric compression of the marginal profile remains near-marginal and benefits from enhanced stability. As the compressed profile is close to the static marginal profile, it is likely undergoing slow  $m = 0$  relaxation on the order of  $1 - 10 \mu\text{s}$  (with an indeterminate role played by flow shear), but voltage decreases on a faster time-scale. We point the reader to [55] for an in-depth discussion of the simulation’s macroevolution.

3) *Observed Temperature Profiles:* The core plasma temperature and low- $\beta$  edge plasma temperature are observed to be super-adiabatic, meaning that temperatures drop off more slowly than if specific entropy were radially constant. We expect that with higher fidelity physics entropy mode activity would relax temperature to the adiabatic profile. While two-fluid and 3-D physics are expected to change the nature of the marginal state attained in modeling, it seems likely that this conceptual picture of a core plasma with relaxed gradients surrounded by low- $\beta$  sheared flow will continue to hold under appropriate conditions, in line with the stepped relaxation region approach used with success in MHD modeling [59]. We expect three-dimensionality to complicate this picture in surprising ways, for example, the saturated state attained by weakly unstable kink modes in a sheared-flow-stabilized Z-pinch with conducting wall boundary.

## V. CONCLUSION

This work revisited the Kadomtsev profile describing axisymmetric marginal stability of the Z-pinch and investigated its thermodynamic properties, filling a gap in the literature for the static profile. Self-organization based on thermodynamic principles in the magnetic confinement and plasma turbulence communities is a fruitful topic of research [38], [59], [60], [61], so it is hoped that this article contributes to further research in this field by elucidating the properties of self-organization in one of its simplest solvable manifestations other than the force-free Taylor states, namely, the axisymmetrically constrained Z-pinch. In this simple model, we combine the well-known result from the minimum energy principle together with the thermodynamics of a perfect gas to demonstrate that the state marginal to both the interchange and drift-entropy modes is one where the gradients of both specific entropy and magnetic flux are zero. This result is happily in accordance with the equivalence between minimum energy and maximum entropy principles and establishes a clear analogy to the Schwarzschild–Ledoux stability criterion used in other branches of hydrodynamics. However, we emphasize that the core result, namely, Kadomtsev’s stability function,

is most easily determined by Kadomtsev's minimum energy method. We do point out, however, that an interesting avenue for future research is to deduce equilibrium stability as a condition on  $s_z$  using an entropy functional method.

In addition, the analogy between Kadomtsev's pinch and the adiabatic profile used in meteorology and oceanography was used to draw out parallels to sheared-flow stabilization of the respective interchange modes in these fields, and how these marginal states are closely related to statistically probable distributions, here referred to as frozen Boltzmann distributions. Indeed, Kadomtsev's pinch is precisely this sort of frozen Boltzmann distribution in the magnetic energy. This was then extended to a demonstration in Section III-G that the pinch radius itself also follows an exponential relationship in the ratio of magnetic to thermal energies. It was also observed in Section III-H that the extensive total enthalpy  $H$  of the Kadomtsev pinch scales with magnetic flux per unit length  $\psi'$  as  $H = I_\infty \psi'$  for all the values of  $\gamma$ , with the significance that unstable profiles like the sharp pinch can be said to contain an excess "unmixed" thermal enthalpy.

Furthermore, theory is presented here suggesting that non-isothermal temperature profiles are more likely than isothermal ones in the collisional fluid regime according to the activity of the entropy modes. This conclusion is further supported by the results of Section IV where it was shown that the specific enthalpy of an isothermal Kadomtsev-stable pinch diverges as  $\beta \rightarrow 0$  in the edge plasma, but the specific enthalpy is constant for adiabatic (i.e., constant specific entropy) profiles. Indeed, the simulation results of the FuZE experiment shown in Section IV-A demonstrate nonisothermal profiles even without two-fluid physics. These simulation results indicate attraction to the marginal state in the flow pinch formed downstream from a coaxial accelerator, and the sustainment of a near-marginal state following pinch "zippering"/compression.

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